

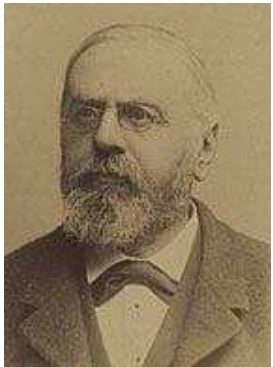
# Der Sinus an der Wiege der Naturwissenschaften

G. Wanner (“Webinar” Genf  $\Rightarrow$  Tübingen, Feb. 2021)

“Wiege ??” ... das beginnt mit etwas ganz Altem ...



A.H. Rhind  
(1833–1863)



A. Eisenlohr



Rhind Papyrus, 1950 v.Chr.  $\Rightarrow$  1650 v.Chr..

**A.H.Rhind** ⇒ August Eisenlohr (1877):

Eisenlohr nr.

Fl. Rechteck 49

Fl. Kreis 50

Fl. Dreieck 51

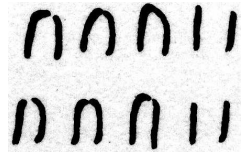
Fl. Trapez 52

Feldteilung 53

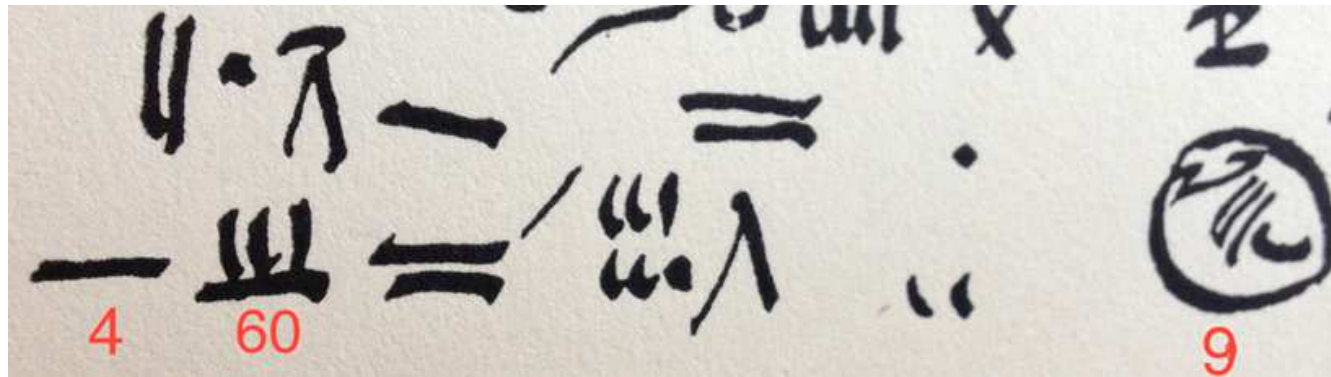


Rhind Nr. 50: **Kreisfl.** =  $(\text{diam} - \frac{\text{diam}}{9})$  **zum Quadrat**

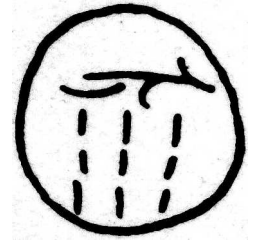
T.E.Peet (1923)



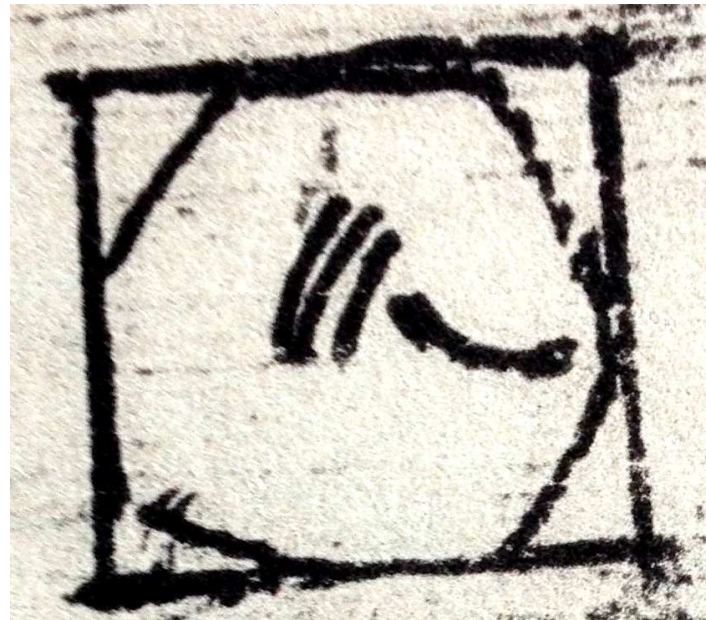
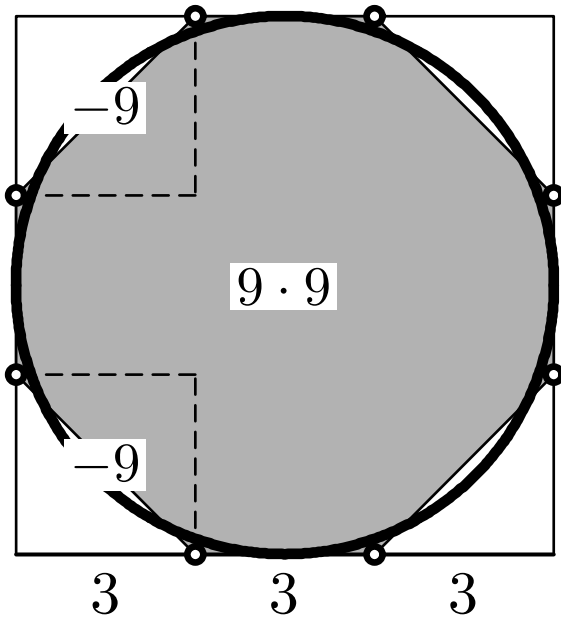
hieroglyphic



hieratic



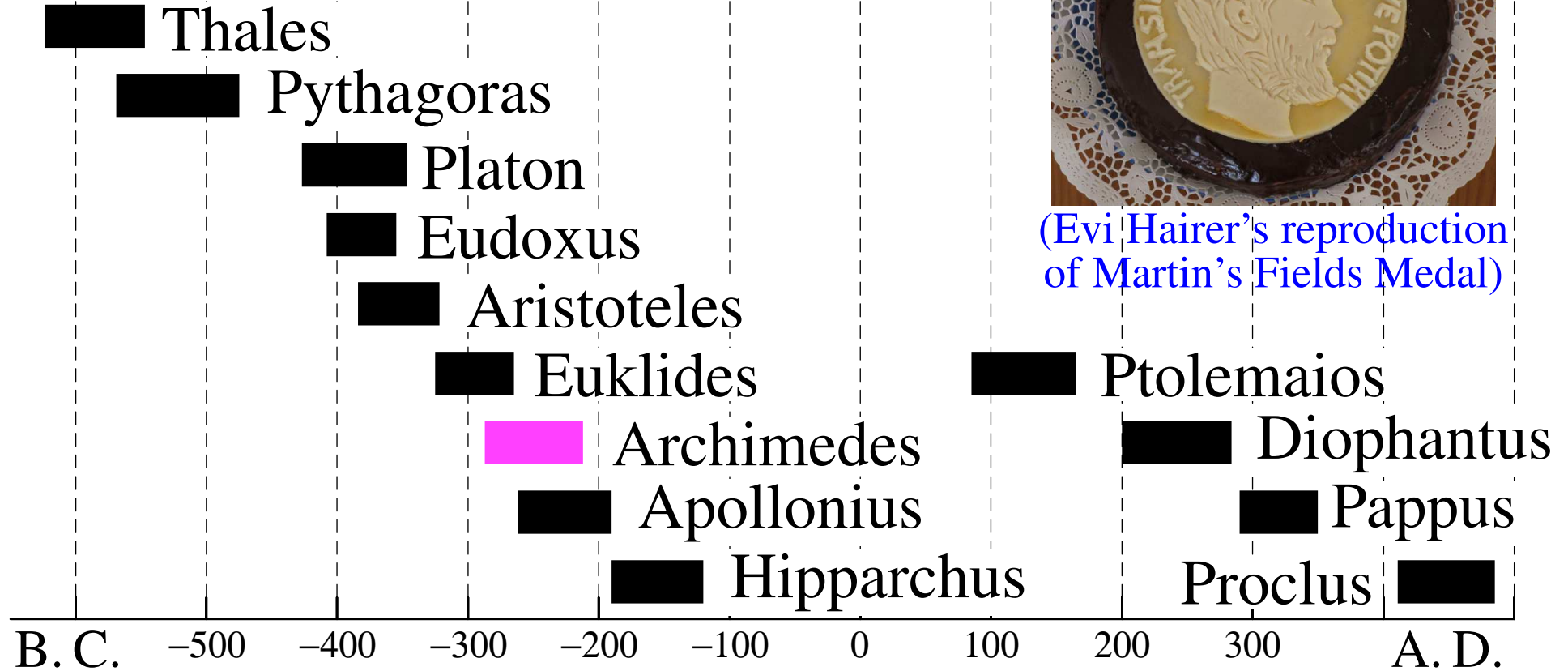
Warum? Rhind Nr. 48:



Fläche =  $9 \cdot 9 - 2 \cdot 9 = 63$  (ist kein Quadrat)  $\approx 64 = 8^2$

(exakt ????? NEIN!)

# Die griechische Strenge: Archimedes



(Evi Hairer's reproduction of Martin's Fields Medal)



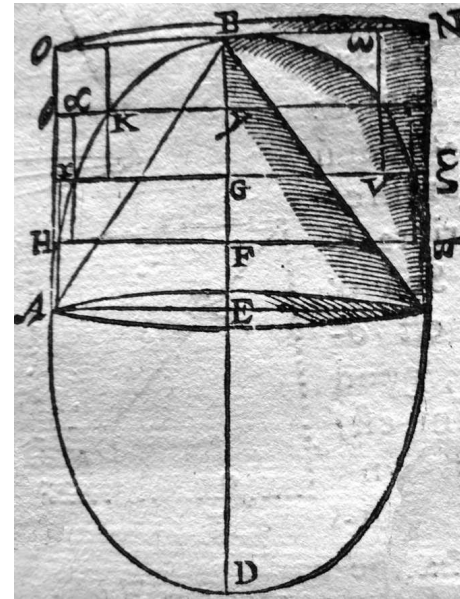
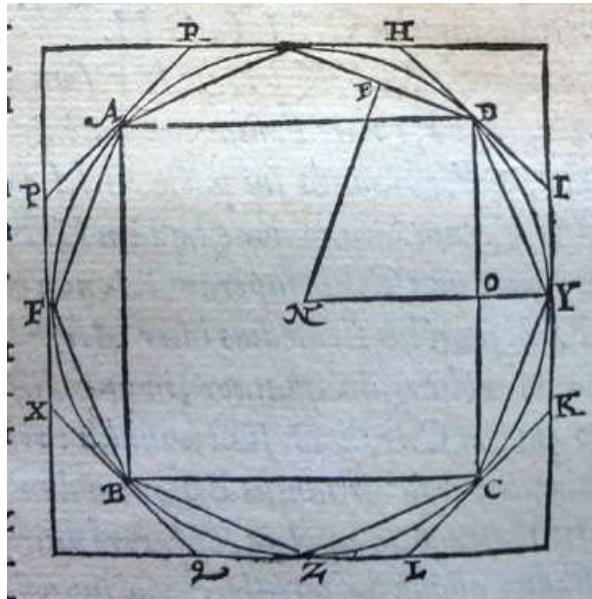
(Archimedes, "Kreismessung", impr. 1665)

Suchen  $\pi = \frac{\text{Perimeter}}{\text{Fl.}}$   = Fl. 

W. Jones 1706

**Euklid** (Buch V): Einzig **streng** ist:  $r < \pi < R$   $r, R$  **rat.**

**Archimedes** “Riemannsche Unter- und Obersummen”:



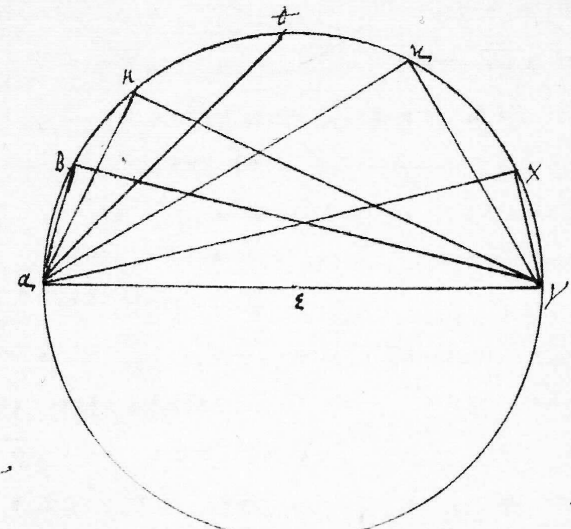
**Proposition 3:**  $(3.1408 <)$   $3\frac{10}{71} < \pi < 3\frac{1}{7}$   $(< 3.1428572)$

**Rhind:**  $\frac{4 \cdot 64}{81} = 3.1605$

# Archimedes' Beweis für $3\frac{10}{71} < \pi$ :

καὶ κέντρον τῆς· καὶ ἡ γὰρ ἀφαιρούμενη· καὶ ἡ  
 ὑπὸ β' αὐ τριήκοντος ὀρθῆς· ἡ εβ' ἀρκ' πρὸς β' γ' λό-  
 γον ἔχει, ὡς τὸ πρὸς ρη'· ἡ δὲ αὐ πρὸς τ'  
 γβ' λόγον ἔχει· ἢ ὅν σξξ' πρὸς ρη'· τατμήσθω οὖν  
 ἡ ὑπὸ β' αὐ δίχα τῆς η'· ἐστὶν ἄρα ὡς ἡ βε' πρὸς  
 αὐ, ἡ βη' πρὸς ηγ'· καὶ ἀναλλάξ καὶ σίμωθάντι·  
 ὡς ἄρα σὺν αμεί' πρὸς ἡ βε' αὐ πρὸς β' γ'· ἡ αὐ  
 πρὸς γη'· ὡς τε ἡ γε' πρὸς γη' μέγιστον λόγον ἔχει·  
 ἢ πρὸς αα' πρὸς ρη'· ἢ αμ' ἄρα πρὸς ηγ' δὴ αμεί'  
 λόγον ἔχει ὅν λδμ' εγν' πρὸς μδ' τυθ' μήκει ἄρα  
 ὅν φηα' πρὸς ρη'· πάλιν δίχα ἡ ὑπὸ ηγ' τῆ εθ'  
 διὰ ταῦτα ἄρα ἡ αὐ πρὸς γθ' μέγιστον λόγον ἔχει  
 ἢ ὅν ρδξβ' ἢ πρὸς ρη'· ἡ εθ' ἄρα πρὸς εγ' μέγισ-  
 τον λόγον ἔχει· ἢ ὅν ρροβη' πρὸς ρη' ἔτι δίχα  
 ἡ ὑπὸ εθ' τῆ ακ' ἡ αὐ ἄρα πρὸς γκ' μέγισ-  
 τον λόγον ἔχει· ἢ ὅν β' τ λθ', ὅν πρὸς ρη'· ἢ εκ'  
 ἄρα πρὸς γκ' μέγιστον· ἢ ὅν β' τ λθ' λ' πρὸς  
 ρη'· ἔτι δίχα ἡ ὑπὸ κεν' τῆ λδ'· ἡ αὐ ἄρα  
 πρὸς λγ' μέγιστον μήκει λόγον ἔχει, ἢ πρὸς ε' μοι  
 πρὸς ρη'· ἐπεὶ οὖν ἡ ὑπὸ β' αὐ τριήκοντος οὐσα  
 ὀρθῆς τατμήσθω τετράκις δίχα, ἡ ὑπὸ καὐ  
 ὀρθῆς ἐστὶ κοί· καὶ ἡ κμ' ἄρα αὐθ' φά του κύκλου  
 κέντρου ἔστι πολυγώνη πλεῖρα ἔχοντος 95·  
 ἐπεὶ οὖν ἡ αὐ πρὸς τὴν γλ' ἀδύχθη μέγιστον λόγον  
 ἔχει ἔχουσα· ἢ πρὸς δ' χοτ' πρὸς ρη'· ἀλλὰ  
 τῆσ' μὲν αὐ δίπλα ἡ αὐ· τῆσ' δὲ γλ' δὴ πλεονῶν  
 ἡ κμ'· καὶ ἡ αχ' αὐ ἄρα πρὸς τὴν τοῦ 95  
 πολυγώνη περίμετρον, μέγιστον λόγον ἔχει· ἢ πρὸς  
 δ' χοι' πρὸς μ' ρ' χ' π'· καὶ ἐστὶ τρίπλασιον καὶ

ελάχιστον λόγον ἔχει, ἢ ὅν β' τ κθ' ἐστὶ πρὸς τ' π'· ἢ  
 ὅν ρ' ω' κγ' πρὸς σν'· ἕκαστα γὰρ ἕκαστα,  
 ἢ γ' α'· ὡς τε ἡ αὐ πρὸς τὴν γθ'· ἢ ὅν ρ' ω' λη' θ'  
 πρὸς σμ'· ἐπεὶ δίχα ἡ ὑπὸ εθ' τῆ κα'· καὶ ἡ αλ'  
 πρὸς τὴν κγ'· ἄρα ελάχιστον λόγον ἔχει, ἢ ὅν ρ' ε'  
 πρὸς εξξ'· ἕκαστα γὰρ ἕκαστα, οἷμαι ἄρα πρὸς  
 τὴν κατ' εἶρον· ρος καὶ πρὸς εξξ'· ἐπεὶ δίχα  
 ἡ ὑπὸ καὐ τῆ λδ'· ἡ αλ' ἄρα πρὸς τὴν αὐ ελίσ-  
 τιστον λόγον ἔχει ἢ ὅν ρ' ε' τ' α' β' ις' ε' πρὸς εξξ'· ἢ δ'  
 εγ' πρὸς γλ' ελάχιστον, ἢ τ' α' β' ις' ε' πρὸς εξξ'· ἀνα-  
 παλιν ἄρα ἐστὶ ἡ περίμετρος τοῦ πολυγώνου πρὸς  
 τὴν διάμετρον μέγιστον λόγον ἔχει ἢ πρὸς ε' τ' α' β' ις' ε'  
 πρὸς εξξ'· ἀπὸ τῶν β' ις' ε', μέγιστον ἐστὶ ἢ τρίπλασι-  
 ον τοῦ ἀντὶ τοῦ κύκλου, τῆσ' δὲ διαμέτρον τρίπλασιον  
 ἐστὶ· καὶ μέγιστον ἢ ὅν δ' ια', ὡς τε καὶ ὁ κύκλος, ἐστὶ  
 κείρον τρίπλασιον ἐστὶ καὶ μέγιστον ἢ δ' ια'· ἢ ἄρα  
 τοῦ κύκλου περίμετρος, τῆσ' διαμέτρον τρίπλασι-  
 ον ἐστὶ· καὶ ελάχιστων μὲν ἢ εβ' δ' ἡ μέρη, μέγιστων δὲ:



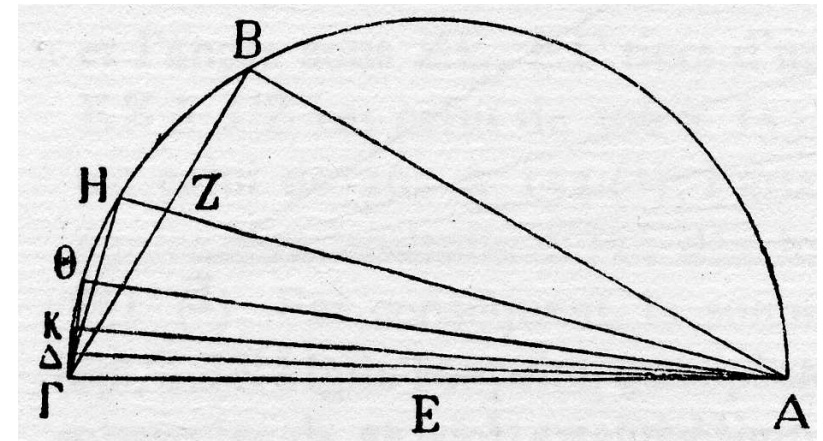
# Entziffern ...

F. Commandino (grec-latin) 1558

Heiberg (grec) 1888

Th. Heath (english) 1920

⇐ Ver Eecke (français) 1921



$$B\Gamma = \frac{1}{2} \text{ (6-gon)}$$

berechne  $H\Gamma$ ,  $\Theta\Gamma$ ,  $K\Gamma$ ,  $\Lambda\Gamma$

und  $96 \cdot \Lambda\Gamma < \pi$ .

d'où :  $\frac{A\Gamma}{A\Gamma+AB} = \frac{\Gamma Z}{\Gamma Z+ZB}$ , ou  $\frac{A\Gamma}{\Gamma Z} = \frac{A\Gamma+AB}{B\Gamma}$ , d'où :  $\frac{A\Gamma}{H\Gamma} = \frac{A\Gamma+AB}{B\Gamma} = \frac{A\Gamma}{B\Gamma} + \frac{AB}{B\Gamma}$ . Or,  
 $\frac{A\Gamma}{B\Gamma} = \frac{1560}{780}$ , et  $\frac{AB}{B\Gamma} < \frac{1351}{780}$ ; donc  $\frac{A\Gamma}{H\Gamma} < \frac{1560+1351}{780}$ , ou  $\frac{A\Gamma}{H\Gamma} < \frac{2911}{780}$ . D'autre part,  
 $\frac{\overline{A\Gamma}^2}{\overline{H\Gamma}^2} < \frac{(2911)^2}{(780)^2}$ , d'où :  $\frac{\overline{A\Gamma}^2 + \overline{H\Gamma}^2}{\overline{H\Gamma}^2} < \frac{2911^2 + 780^2}{780^2}$ , ou  $\frac{\overline{A\Gamma}^2}{\overline{H\Gamma}^2} < \frac{9082321}{608400}$ , d'où,  
 comme le texte :  $\frac{A\Gamma}{H\Gamma} < \frac{3013\frac{3}{4}}{780}$ .

1. On aura, comme dans le cas précédent :  $\frac{A\Theta}{\Theta\Gamma} = \frac{A\Gamma+AH}{H\Gamma} = \frac{A\Gamma}{HA} + \frac{AH}{H\Gamma}$ , d'où, sub-  
 stituant les valeurs de ces deux derniers termes :  $\frac{A\Theta}{\Theta\Gamma} < \frac{3013\frac{3}{4} + 2911}{780}$ , ou  
 $\frac{A\Theta}{\Theta\Gamma} < \frac{5924\frac{3}{4}}{780}$ , ou  $\frac{A\Theta}{\Theta\Gamma} < \frac{\frac{4}{13} \times 5924\frac{3}{4}}{\frac{4}{13} \times 780}$ , ou  $\frac{A\Theta}{\Theta\Gamma} < \frac{1823}{240}$ . D'autre part,  $\frac{\overline{A\Theta}^2}{\overline{\Theta\Gamma}^2} < \frac{1823^2}{240^2}$ ,  
 d'où  $\frac{\overline{A\Theta}^2 + \overline{\Theta\Gamma}^2}{\overline{\Theta\Gamma}^2} < \frac{1823^2 + 240^2}{240^2}$ , ou  $\frac{\overline{A\Gamma}^2}{\overline{\Theta\Gamma}^2} < \frac{3380929}{57600}$ , d'où, comme le texte :  
 $\frac{A\Gamma}{\Theta\Gamma} < \frac{1838\frac{9}{11}}{240}$ .

2. On aura de même :  $\frac{AK}{K\Gamma} = \frac{A\Gamma+AK}{\Theta\Gamma} = \frac{A\Gamma}{\Theta\Gamma} + \frac{AK}{\Theta\Gamma}$ . et, par substitution des valeurs  
 trouvées pour ces deux derniers termes, il vient :  $\frac{AK}{K\Gamma} < \frac{1838\frac{9}{11} + 1823}{240}$ , ou  
 $\frac{AK}{K\Gamma} < \frac{3661\frac{9}{11}}{240}$ , ou  $\frac{AK}{K\Gamma} < \frac{\frac{11}{16} \times 3661\frac{9}{11}}{\frac{11}{16} \times 240}$ , ou  $\frac{AK}{K\Gamma} < \frac{1007}{66}$ . D'autre part,  $\frac{\overline{AK}^2}{\overline{K\Gamma}^2} < \frac{1007^2}{66^2}$ ,  
 d'où :  $\frac{\overline{AK}^2 + \overline{K\Gamma}^2}{\overline{K\Gamma}^2} < \frac{1007^2 + 66^2}{66^2}$ , ou  $\frac{\overline{A\Gamma}^2}{\overline{K\Gamma}^2} < \frac{1018405}{14356}$ , d'où, comme le texte :  
 $\frac{A\Gamma}{K\Gamma} < \frac{1009\frac{1}{8}}{66}$ .

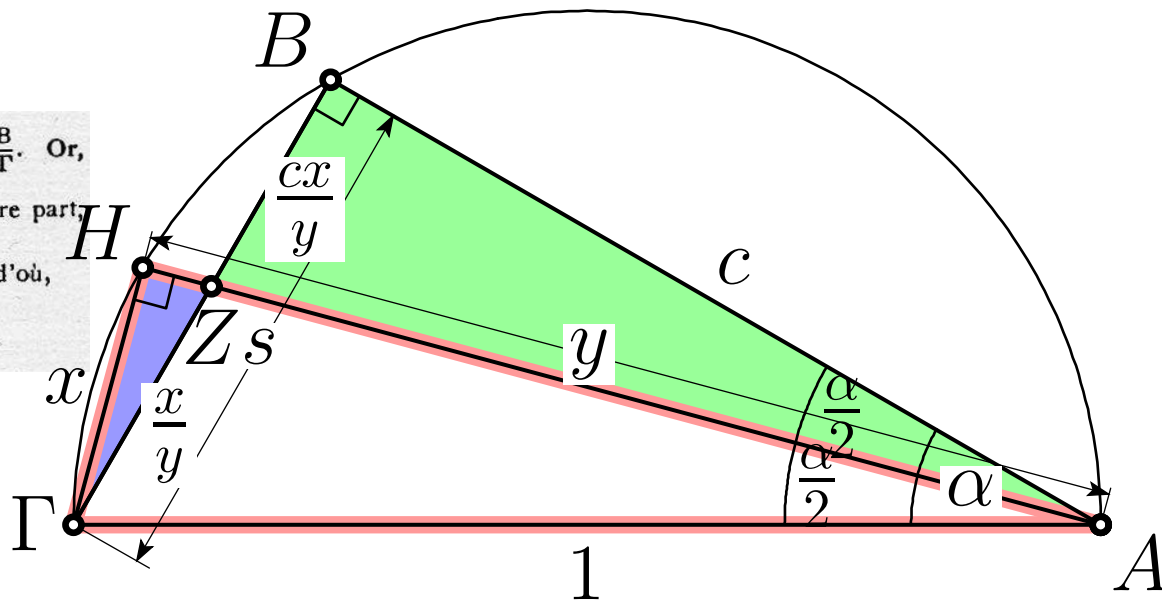
3. On aura de même :  $\frac{A\Lambda}{\Lambda\Gamma} = \frac{A\Gamma+AK}{K\Gamma} = \frac{A\Gamma}{K\Gamma} + \frac{AK}{K\Gamma}$ , et, par substitution des valeurs  
 précédentes :  $\frac{A\Lambda}{\Lambda\Gamma} < \frac{1009\frac{1}{8}}{66} + \frac{1007}{66}$ , ou  $\frac{A\Lambda}{\Lambda\Gamma} < \frac{2016\frac{1}{8}}{66}$ . D'autre part,  
 $\frac{\overline{A\Lambda}^2 + \overline{\Lambda\Gamma}^2}{\overline{\Lambda\Gamma}^2} < \frac{(2016\frac{1}{8})^2 + 66^2}{66^2}$ , ou  $\frac{\overline{A\Gamma}^2}{\overline{\Lambda\Gamma}^2} < \frac{4069284\frac{1}{8}}{4356}$ , d'où, comme le texte :  $\frac{A\Gamma}{\Lambda\Gamma} < \frac{2017\frac{1}{4}}{66}$ .

1. Sous-entendu : περίμετρος, le périmètre (du cercle).  
 2. La relation de la note avant-précédente donne, par inversion :  $\frac{A\Gamma}{\Lambda\Gamma} > \frac{66}{2017\frac{1}{4}}$ ,  
 d'où, observant que  $96 \times \Lambda\Gamma =$  périmètre polygone inscrit de 96 côtés :  
 $\frac{\text{périmètre polygone de 96 côtés}}{\text{diamètre cercle}} > \frac{96 \times 66}{2017\frac{1}{4}}$ , ou  $> \frac{6336}{2017\frac{1}{4}}$ . Or,  $\frac{6336}{2017\frac{1}{4}} > 3\frac{10}{11}$ ,  
 d'où périmètre polygone de 96 côtés  $> 3\frac{10}{11}$  diamètre cercle, d'où, à fortiori, suivant  
 le texte : Circonférence cercle  $> 3\frac{10}{11}$  diamètre.

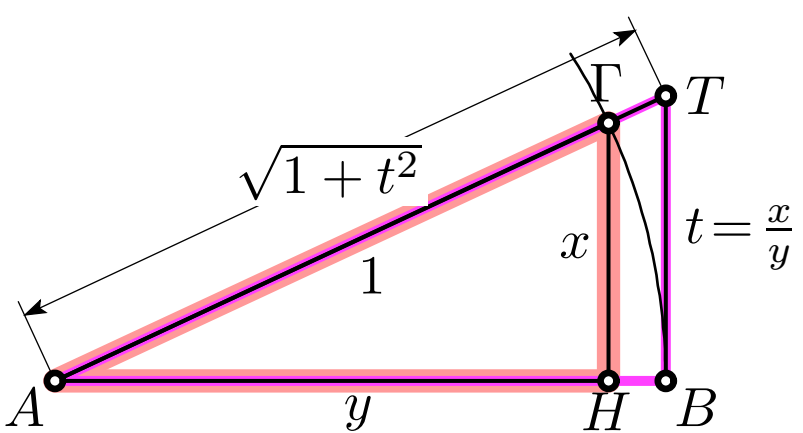
Noch etwas klarer:  $(s, c)_{\text{pour } \alpha} \mapsto (x, y)_{\text{pour } \frac{\alpha}{2}}$  :

d'où :  $\frac{A\Gamma}{A\Gamma+AB} = \frac{\Gamma Z}{\Gamma Z+ZB}$ , ou  $\frac{A\Gamma}{\Gamma Z} = \frac{A\Gamma+AB}{B\Gamma}$ , d'où :  $\frac{A\Gamma}{H\Gamma} = \frac{A\Gamma+AB}{B\Gamma} = \frac{A\Gamma}{B\Gamma} + \frac{AB}{B\Gamma}$ . Or,  
 $\frac{A\Gamma}{B\Gamma} = \frac{1560}{780}$ , et  $\frac{AB}{B\Gamma} < \frac{1351}{780}$  ; donc  $\frac{A\Gamma}{H\Gamma} < \frac{1560+1351}{780}$ , ou  $\frac{A\Gamma}{H\Gamma} < \frac{2911}{780}$ . D'autre part,  
 $\frac{A\Gamma^2}{H\Gamma^2} < \frac{(2911)^2}{(780)^2}$ , d'où :  $\frac{A\Gamma^2 + H\Gamma^2}{H\Gamma^2} < \frac{2911^2 + 780^2}{780^2}$ , ou  $\frac{A\Gamma^2}{H\Gamma^2} < \frac{9082321}{608400}$ , d'où,  
 comme le texte :  $\frac{A\Gamma}{H\Gamma} < \frac{3013\frac{3}{4}}{780}$ .

(Ver Ecke)



$$\triangle \approx \triangle \Rightarrow \Gamma Z = \frac{x}{y}, \quad \triangle \approx \triangle \Rightarrow ZB = \frac{cx}{y} \Rightarrow s = \frac{x}{y}(1 + c)$$



$$\frac{x}{y} = \frac{s}{c+1} = t \quad \text{sin. } \frac{1}{2} v = \sqrt{\frac{1 - \text{cos. } v}{2}}$$

$$\triangle \approx \triangle \Rightarrow x = \frac{t}{\sqrt{1+t^2}} = \sqrt{\frac{1-c}{2}}$$

$$H\Gamma > \frac{780}{3013\frac{3}{4}}, \quad \Theta\Gamma > \frac{240}{1838\frac{9}{11}}, \quad K\Gamma > \frac{66}{1009\frac{1}{6}}, \quad \Lambda\Gamma > \frac{66}{2017\frac{1}{4}} \Rightarrow \boxed{3\frac{1137}{8069}}$$



... und weiter ?

Zu Chongzhi (China 480 n.Chr.) mit derselben Methode:

$$3.1415926 < \pi < 3.1415927$$

Mādhava (Kerala, Indien 1350–1425 n.Chr.) ähnliche M.:

$$\pi = 3.14159265359$$

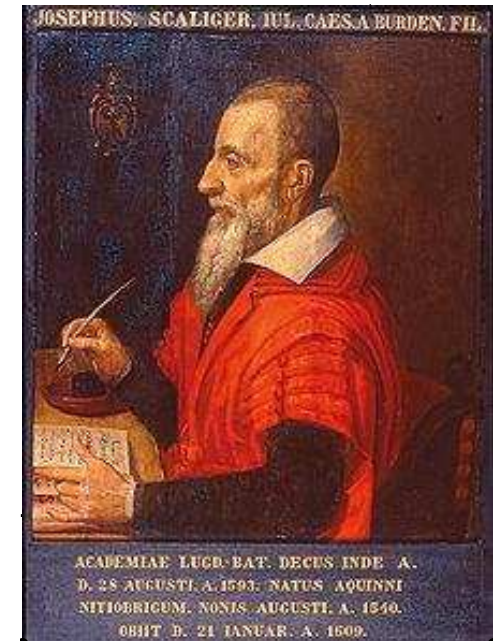
Al-Kāshī (Samarkand 1424 n.Chr.)  $26 \times$  Archimedes :

$$\pi = 3.1415926535897932$$

Die Europäer waren noch etwas hinten:

Josephus Scaliger (Leyden publ. 1594):

$$\pi = \sqrt{10} = 3.162278 \quad !!!$$



Diese Behauptung des berühmten Professors Joseph Scaliger entfachte einen Sturm der Entrüstung ...

IN  
**ARCHIMEDIS**

CIRCVLI DIMENSIONEM  
Expositio & Analysis.

**APOLOGIA PRO ARCHIMEDE**  
ad Clariss. virum Iosephum Scaligerum.

**EXERCITATIONES CYCLICAE**  
contra Iosephum Scaligerum, Orontium Finæum, & Raymarum  
Vrsum, in decem Dialogos distinctæ.

**AVTHORE ADRIANO ROMANO EQVITE**  
AVTARO Mathematicæ Excellentissimo Professore in  
Academia Vrceburgensi.



**WVRCEBURGI**  
ANNO MDCLXCVII

**Adriaan van Roomen**

(publ. WVRCEBURGI 1597):

**nein, Archimedes ist richtig ..**

$$\pi = 3.1415926535897932$$

Sein Freund

**Ludolph van Ceulen**

am Ende seines Lebens (1610)

(Grabinschrift in der

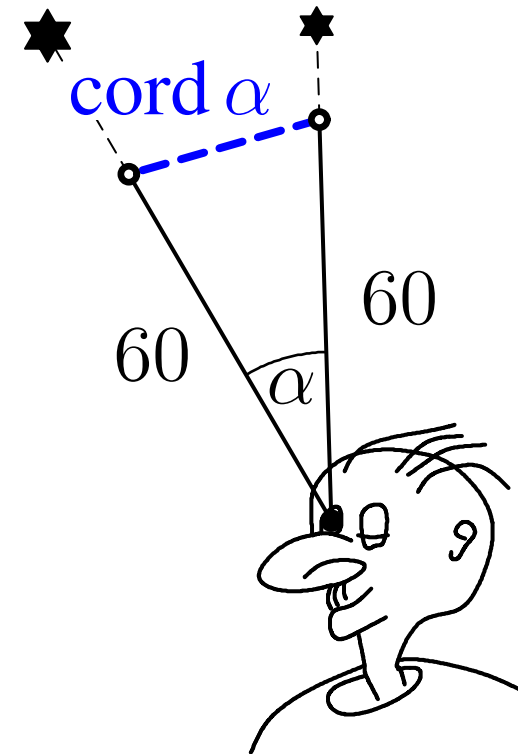
Pieterskerk zu Leyden)...

$$\pi = 3.14159265358979323846264338327950288$$

# Sehntafel des Klaudios Ptolemaios

(150 A.D., Cl. Ptolémée, Ptolemy, Tolomeo, Ptolemæus, Ptolemäus,...)

ΚΑΝΟΝΙΟΝ ΤΩΝ ΕΝ ΚΥΚΛΩ ΕΥΘΕΙΩΝ.									
ΠΕΡΙΦΕΡΕΙΩΝ.		ΕΥΘΕΙΩΝ.			ΕΞΗΚΟΤΣΩΝ.				
Μοιρῶν.		Μ.	Π.	Δ.	Μ.	Π.	Δ.	Τ.	
ὀ	ς"	ὀ	λα	κε	ὀ	α	β	ν	
α	ο'	α	β	ν	ὀ	α	β	ν	
α	ς"	α	λθ	εε	ὀ	α	β	ν	
β	ὀ	β	ε	μ	ὀ	α	β	ν	
β	ς"	β	λς	θ	ὀ	α	β	μη	
γ	ὀ	γ	η	κη	ὀ	α	β	μη	
γ	ς"	γ	λθ	νβ	ὀ	α	β	μη	
δ	ο'	δ	ια	ις	ὀ	α	β	μς	
δ	ς"	δ	μβ	μ	ὀ	α	β	μς	
ε	ὀ	ε	ιθ	θ	ὀ	α	β	μς	
ε	ς"	ε	με	κς	ὀ	α	β	με	
ς	ο'	ς	ις	μθ	ὀ	α	β	μθ	
ς	ς"	ς	μη	ια	ὀ	α	β	μγ	
ς	ο'	ς	ιθ	λγ	ὀ	α	β	μβ	
ς	ς"	ς	ν	νθ	ὀ	α	β	μα	
η	ὀ	η	κβ	εε	ὀ	α	β	μ	
ς	ς"	η	νγ	λε	ὀ	α	β	λθ	
θ	ο'	θ	κθ	νθ	ὀ	α	β	λη	



# Wie zu lesen ?

ΚΑΝΟΝΙΟΝ ΤΩΝ ΕΝ ΚΥΚΛΩ ΕΥΘΕΙΩΝ.								
ΠΕΡΙΦΕΡΕΙΩΝ.		ΕΥΘΕΙΩΝ.			ΕΞΗΚΟΤΩΝ.			
Μοιρῶν.		Μ.	Π.	Δ.	Μ.	Π.	Δ.	Τ.
ο̄	ς̄"	ο̄	λα	κε	ο̄	α	β	ν
α	ο̄"	α	β	ν	ο̄	α	β	ν
α	ς̄"	α	λθ	εε	ο̄	α	β	ν
β	ο̄"	β	ε	μ	ο̄	α	β	ν
β	ς̄"	β	λς	δ	ο̄	α	β	μη
γ	ο̄"	γ	η	κη	ο̄	α	β	μη
γ	ς̄"	γ	λθ	νβ	ο̄	α	β	μη
δ	ο̄"	δ	ια	ις	ο̄	α	β	μς
δ	ς̄"	δ	μβ	μ	ο̄	α	β	μς
ε	ο̄"	ε	ιθ	δ	ο̄	α	β	μς
ε	ς̄"	ε	με	κς	ο̄	α	β	με
ς	ο̄"	ς	ις	μθ	ο̄	α	β	μδ
ς	ς̄"	ς	μη	ια	ο̄	α	β	μγ
ς	ο̄"	ς	ιθ	λγ	ο̄	α	β	μβ
ς	ς̄"	ς	ν	νδ	ο̄	α	β	μα
η	ο̄"	η	κβ	εε	ο̄	α	β	μ
η	ς̄"	η	νγ	λε	ο̄	α	β	λθ
θ	ο̄"	θ	κδ	νδ	ο̄	α	β	λη

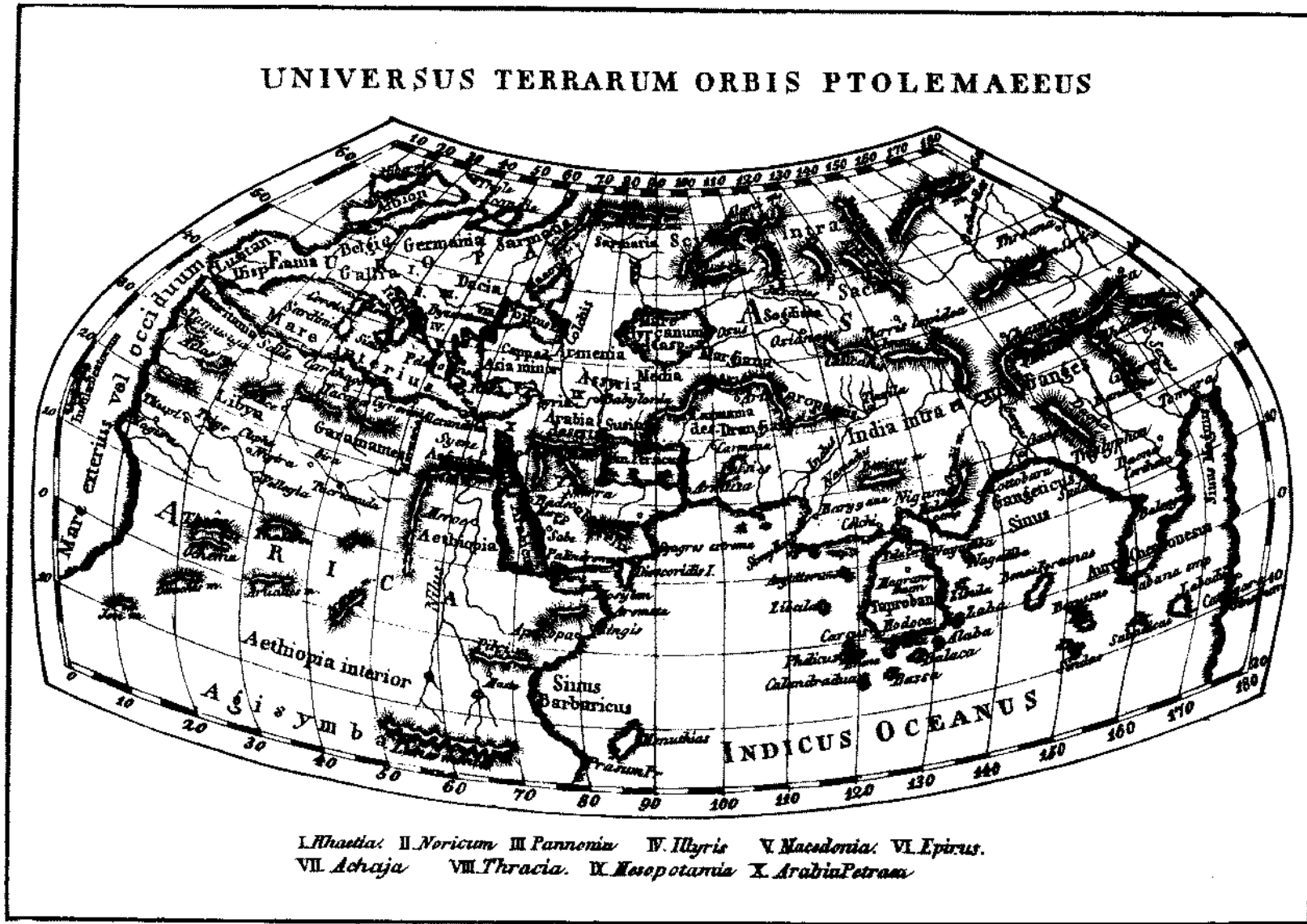
gr.		gr.	rm.	
α	1	A	A	
β	2	B	B	
γ	3	Γ	C	↓
δ	4	Δ	D	
ε	5	E	E	
(F)	6	F	F	
ς	7	Z ↗		G
η	8	H	H	
θ	9	Θ ↗		I
ι	10	I	I	J
κ	20	K	K	
λ	30	Λ	L	
μ	40	M	M	Z
ν	50	N	N	↓

zur Kontrolle:

ΚΑΝΟΝΙΟΝ ΤΩΝ ΕΝ ΚΥΚΛΩ ΕΥΘΕΙΩΝ.								
ΠΕΡΙΦΕ- ΡΕΙΩΝ.		ΕΥΘΕΙΩΝ.			ΕΞΗΚΟΤΣΩΝ.			
Μοιρῶν.		Μ.	Π.	Δ.	Μ.	Π.	Δ.	Τ.
ο̄	ς̄"	ο̄	λα	κε	ο̄	α	β	ν
α	ο̄"	α	β	ν	ο̄	α	β	ν
α	ς̄"	α	λθ	εε	ο̄	α	β	ν
β	ο̄"	β	ε	μ	ο̄	α	β	ν
β	ς̄"	β	λς	δ	ο̄	α	β	μη
γ	ο̄"	γ	η	κη	ο̄	α	β	μη
γ	ς̄"	γ	λθ	νβ	ο̄	α	β	μη
δ	ο̄"	δ	ια	ις	ο̄	α	β	μς
δ	ς̄"	δ	μβ	μ	ο̄	α	β	μς
ε	ο̄"	ε	ιθ	δ	ο̄	α	β	μς
ε	ς̄"	ε	με	κς	ο̄	α	β	με
ς	ο̄"	ς	ις	μθ	ο̄	α	β	μδ
ς	ς̄"	ς	μη	ια	ο̄	α	β	μγ
ς	ο̄"	ς	ιθ	λγ	ο̄	α	β	μβ
ς	ς̄"	ς	ν	νδ	ο̄	α	β	μα
η	ο̄"	η	κβ	εε	ο̄	α	β	μ
η	ς̄"	η	νγ	λε	ο̄	α	β	λθ
θ	ο̄"	θ	κδ	νδ	ο̄	α	β	λη

0	30	0	31	25
1	0	1	2	50
1	30	1	34	15
2	0	2	5	39
2	30	2	37	4
3	0	3	8	28
3	30	3	39	53
4	0	4	11	17
4	30	4	42	40
5	0	5	14	4
5	30	5	45	27
6	0	6	16	49
6	30	6	48	11
7	0	7	19	33
7	30	7	50	54
8	0	8	22	15
8	30	8	53	35
9	0	9	24	54

# Ptolemaios' Geographie (Γῆ ... γράφω ... μετρέω):



Ptolemaios vermisst ὑφ' οὓς Παρίσιοι, καὶ πόλις  
8000 Städte, z.B. ... Παρισίων Λουκοτεκία . . . . . κ̄η Ϸ μ̄η Ϸ.

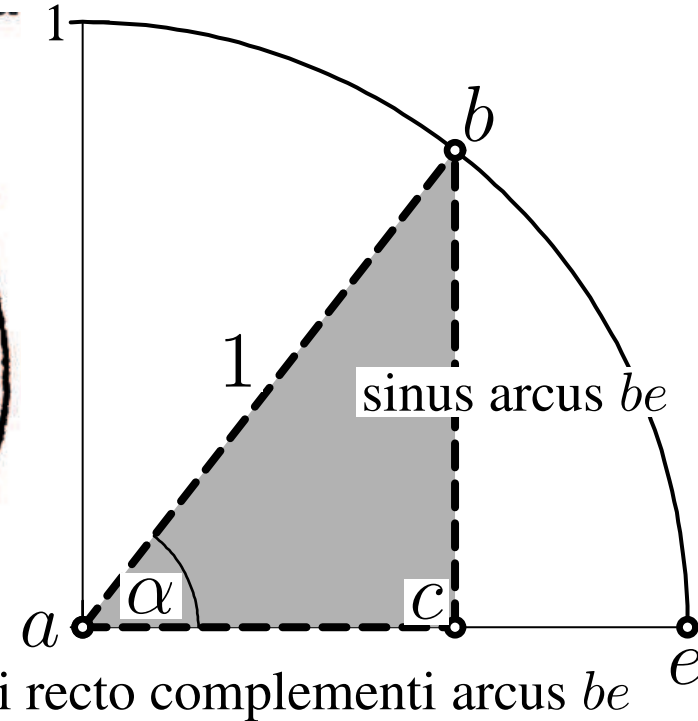
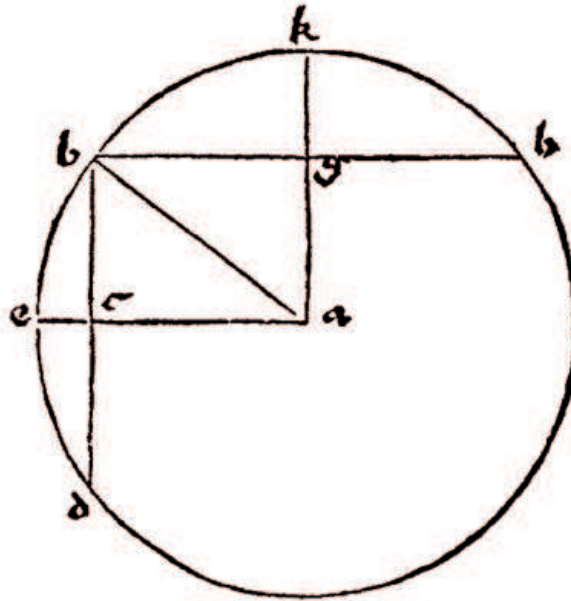
Längengrade fehlerhaft bei 180° ⇒ Chr. Columbus !!

# Johannes Regiomontanus (1436–1476):

mit 11 Uni Leipzig, mit 12 erstes Buch, mit 14 Uni Wien, mit 25 Rom  
liest Archimedes, Ptolemeus im Original, sowie arabische Schriften  $\Rightarrow$

## *De triangulis omnimodis libri quinque*, (publ. 1533);

quoad satis est prolongatū in e puncto. Dico quòd latus b c angulo b a c oppositum est sinus arcus b e dictum angulum subtendens. Latus autem tertium, scilicet a c, æquale est sinui recto complementi arcus b e. Extendatur enim latus b c occurrendo circumferentiæ circuli in puncto d. à punctis autem a quidem centro circuli exeat semidiameter a k æquedistans lateri b c. & à puncto b corda b h æquedistans lateri a c. secabunt autem se necessario duæ lineæ b h & a k, angulis a b h & b a k acutis existentibus, quod fiat in puncto g. Quia itaq; semidi-



## *Compositio tabularum sinuum* (publ. 1541)

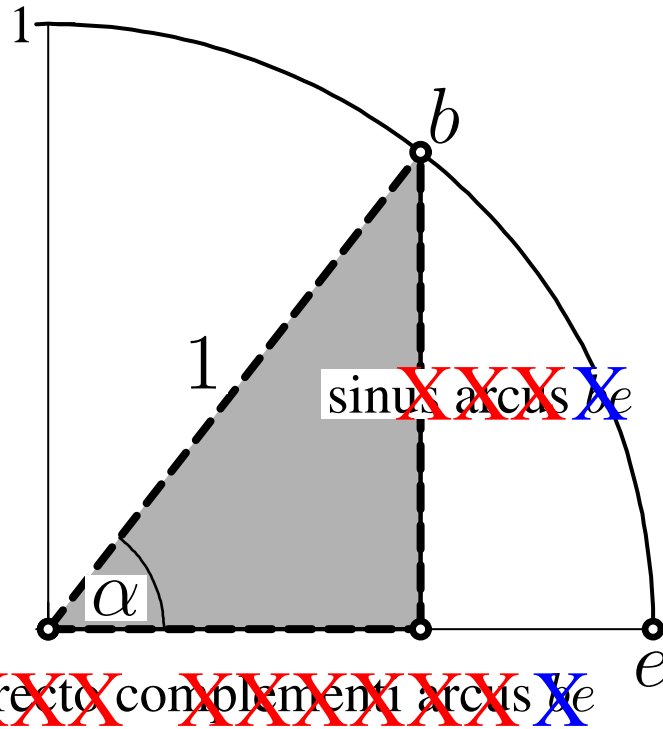
(“SEQVITVR NVNC EIVSDEM IOANNIS Regiomontani tabula sinuum,  
per singula minuta extensa . . .”).



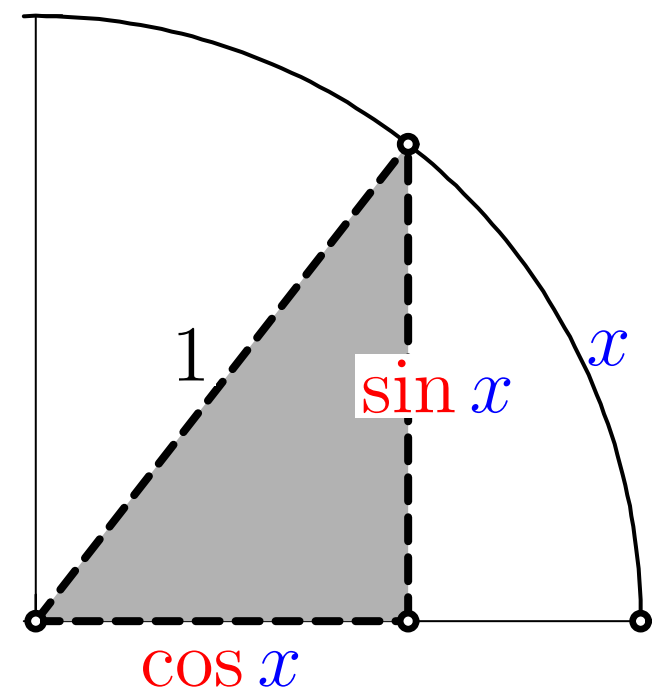
**L. Euler**

**E14** (1729):

**E101** (1748):



⇒



Für die Additionstheoreme (*Introductio* Cap. VIII, §128)

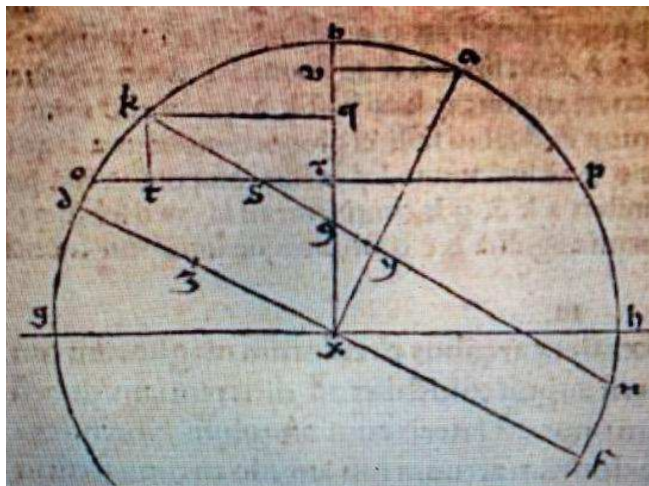
$$\sin(x + y) = \sin x \cos y + \cos x \sin y$$

$$\cos(x + y) = \cos x \cos y - \sin x \sin y$$

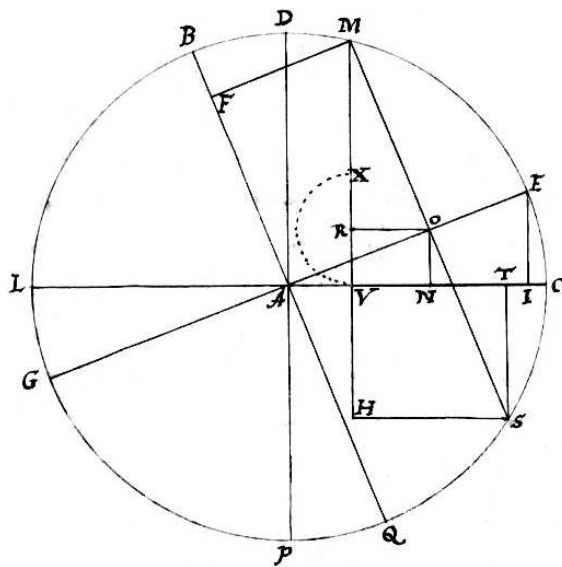
schreibt er einfach **“Hinc vero etiam constat ...”**  
 sie waren also schon vor 3 Jhten. jedem Anfänger bekannt...

**Historische Beweise ??**

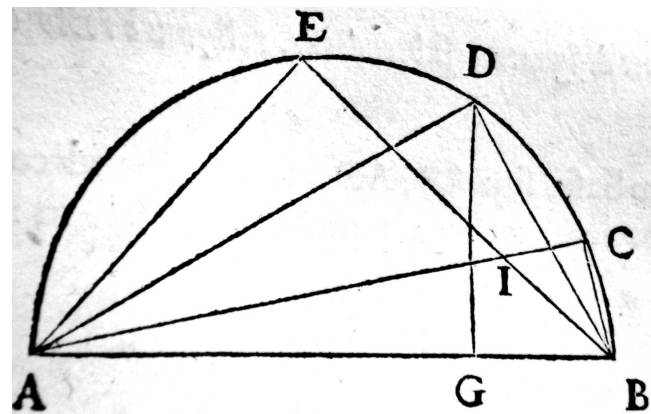




(Regiomontanus 1533, p. 128)

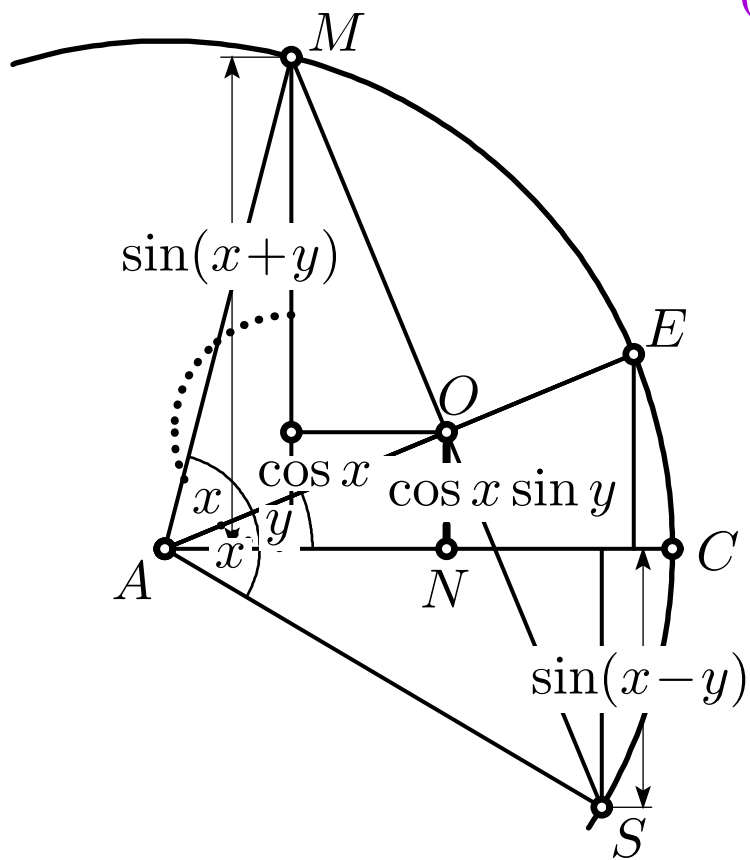


(Jost Bürgi, 1584)

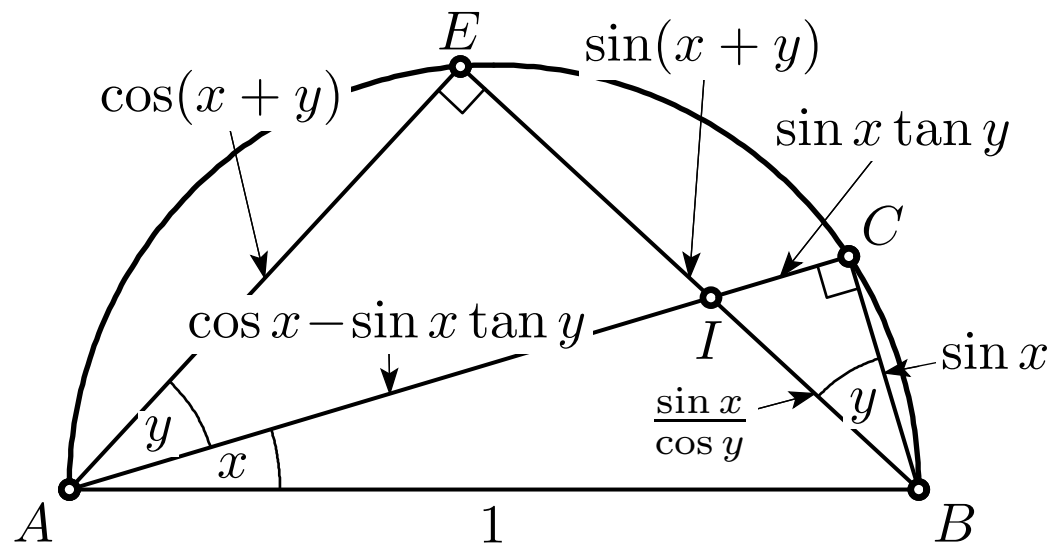


(F. Viète, 1593, impr. 1646)

(vide Archimedes !!)



$$\sin(x + y) - \sin(x - y) = 2 \cos x \sin y$$



$$AE = AI \cos y = \cos x \cos y - \sin x \sin y$$

$$EI = AI \sin y \Rightarrow EI + IB = \sin(x + y)$$

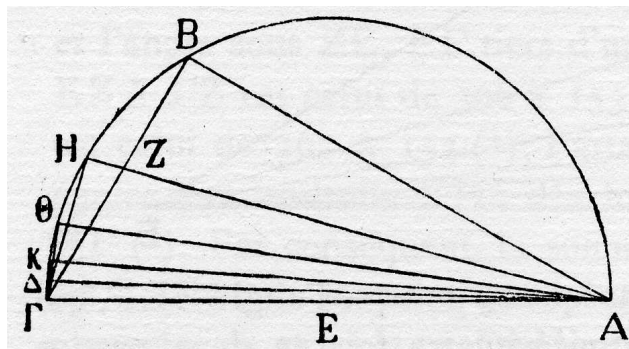
# Historische Berechnungen der Sinustafeln

Das erste Problem war die Berechnung von

$$\sin 1^\circ = 0.01745240643728351 = 1\ 2\ 49\ 43\ 11\ 14\ 44\ 16..$$

alle anderen Sinuswerte durch Additionstheorem.

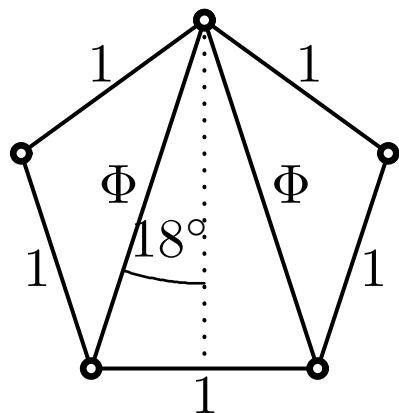
Bereits Archimedes' Resultate für  $B\Gamma$ ,  $H\Gamma$ ,  $\Theta\Gamma$ ,  $K\Gamma$ , ... sind



$n$	$\alpha$	$\sin \alpha$	$\sin \alpha / \alpha$	Basis 60			
6	30.0000	0.50000000	0.01666667	1	0	0	0
12	15.0000	0.25881905	0.01725460	1	2	6	59
24	7.5000	0.13052619	0.01740349	1	2	39	9
48	3.7500	0.06540313	0.01744083	1	2	47	13
96	1.8750	0.03271908	0.01745018	1	2	49	14
192	0.9375	0.01636173	0.01745251	1	2	49	44

Sinuswerte für immer kleinere  $\alpha$ , aber nie  $\alpha = 1^\circ$ .

Lösung:  $\sin 1 \approx \sin \alpha / \alpha$  (bereits genügend für Ptolemäus).



## Bemerkung.

$$\sin 18^\circ = 1/(2\Phi) \text{ (gold. Schnitt)}$$

$$\Rightarrow \sin 3^\circ = \sin(18^\circ - 15^\circ) \text{ (exakt bekannt)}$$

$$x = \sin 1^\circ \Rightarrow -4x^3 + 3x = \sin 3^\circ \text{ (Al-Kāshī 1429)}$$

# Parameśvara (Indien, ca. 1380–1460) (Dank an Ph. Henry für Zitat)

Ist einmal  $\pi$  bekannt (siehe **Mādhava** oben), gibt es einfache Methode für  $\sin \alpha$ ,  $\alpha$  beliebig, beschrieben in Form von 37 Sanskrit Verszeilen, beginnend mit

cāpāj jīvāpi sādhyā syād aviṣeṣākhyakarmanā  
ardhīkuryād iṣṭacāpaṃ tadardham ca tathā punaḥ

die wir aber nicht im Original weiterlesen, auch nicht in der engl. Übersetzung von Kim Plofker (1996), denn aus 37 Verszeilen machen wir 10 Fortranzeilen:

PI=3.14159265358979324D0	0	0.00001704423098	0.99999999985475
S=PI/(1024.D0*180.D0)	1	0.00003408846195	0.99999999941899
C=SQRT(1.D0-S**2)	2	0.00006817692386	0.999999999767595
WRITE(6,*)0,S,C	3	0.00013635384740	0.999999999070381
DO I=1,10	4	0.00027270769226	0.999999996281526
STORE=2.D0*S*C	5	0.00054541536423	0.99999985126103
C=1.D0-2.D0*S**2	6	0.00109083056622	0.99999940504416
S=STORE	7	0.00218165983444	0.99999762017735
WRITE(6,*)I,S,C	8	0.00436330928496	0.99999048072073
END DO	9	0.00872653549880	0.99996192306417
	10	0.01745240643813	0.99984769515638

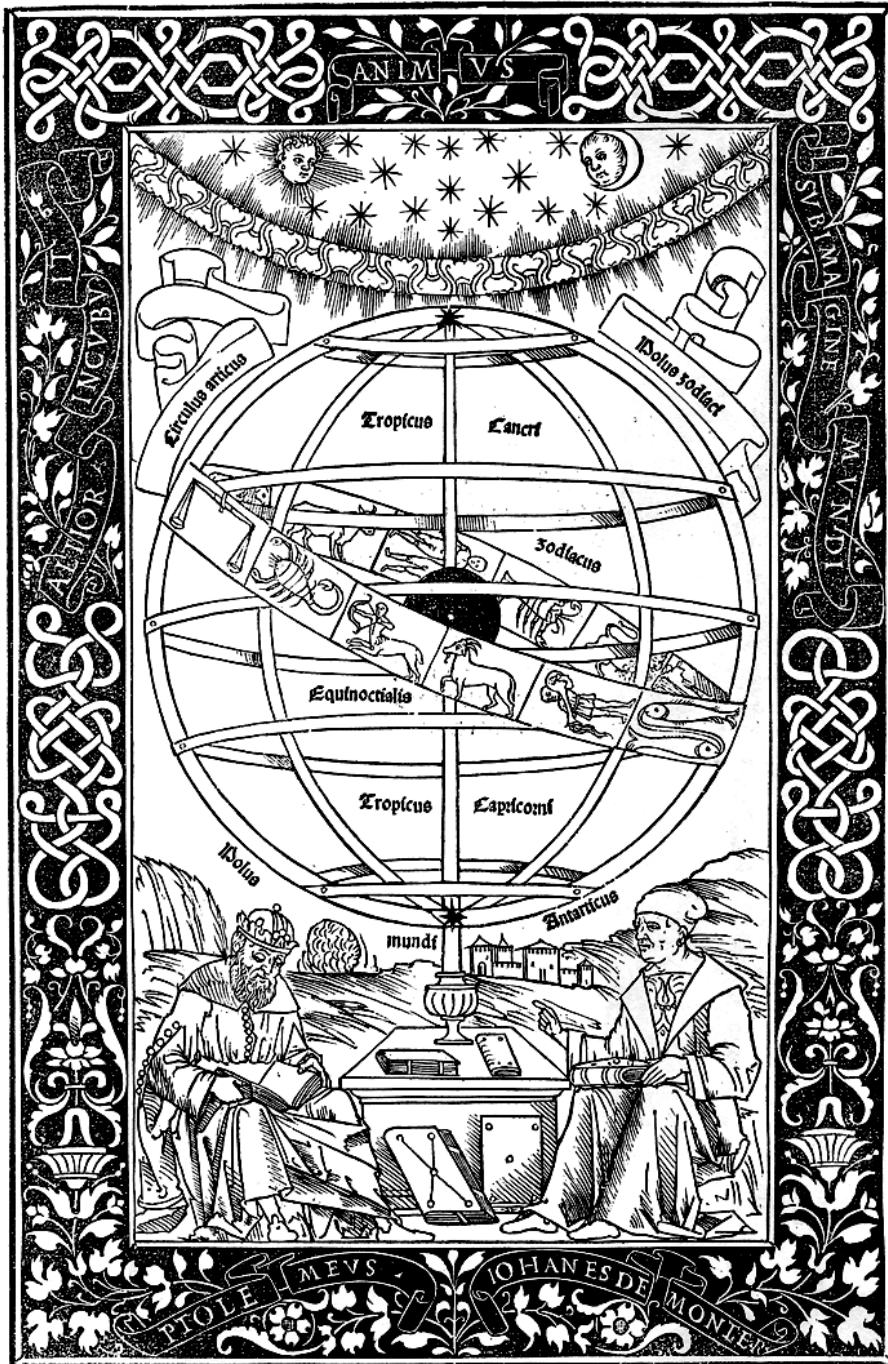
Die **Idee** ist, mit einem beliebigen  $\alpha$  (wir nehmen, what else ?  $\alpha = 1^\circ$ )

dieses  $\alpha$  durch  $2^n$  zu dividieren, mit z.B.  $n = 10$ , dann ist

$\sin(2^{-n}\alpha) \approx \text{arc}(2^{-n}\alpha)$ ,  $\cos = \sqrt{1 - \sin^2}$ , und wir gehen

in  $n$  Schritten mittels  $\sin 2x = 2 \sin x \cos x$ ,  $\cos 2x = 1 - 2 \sin^2 x$  nach oben.

# Ptolemäische Kosmologie



(**Almagest** de Ptolemy, pub. 1496)



(Copernicus, publ. 1543)

# Tycho Brahe (1546 – 1601)

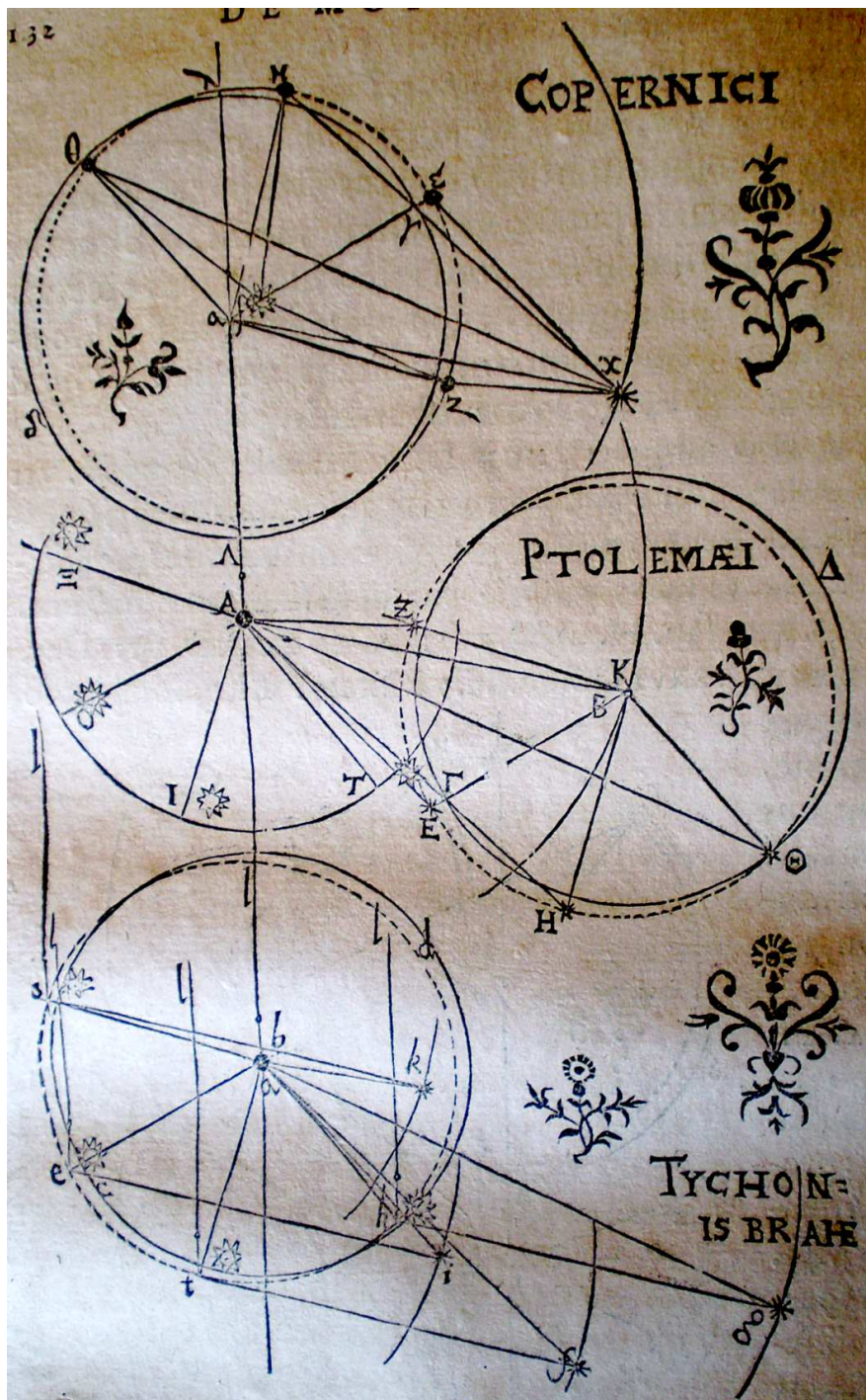
# Joh. Kepler (1571 – 1630)



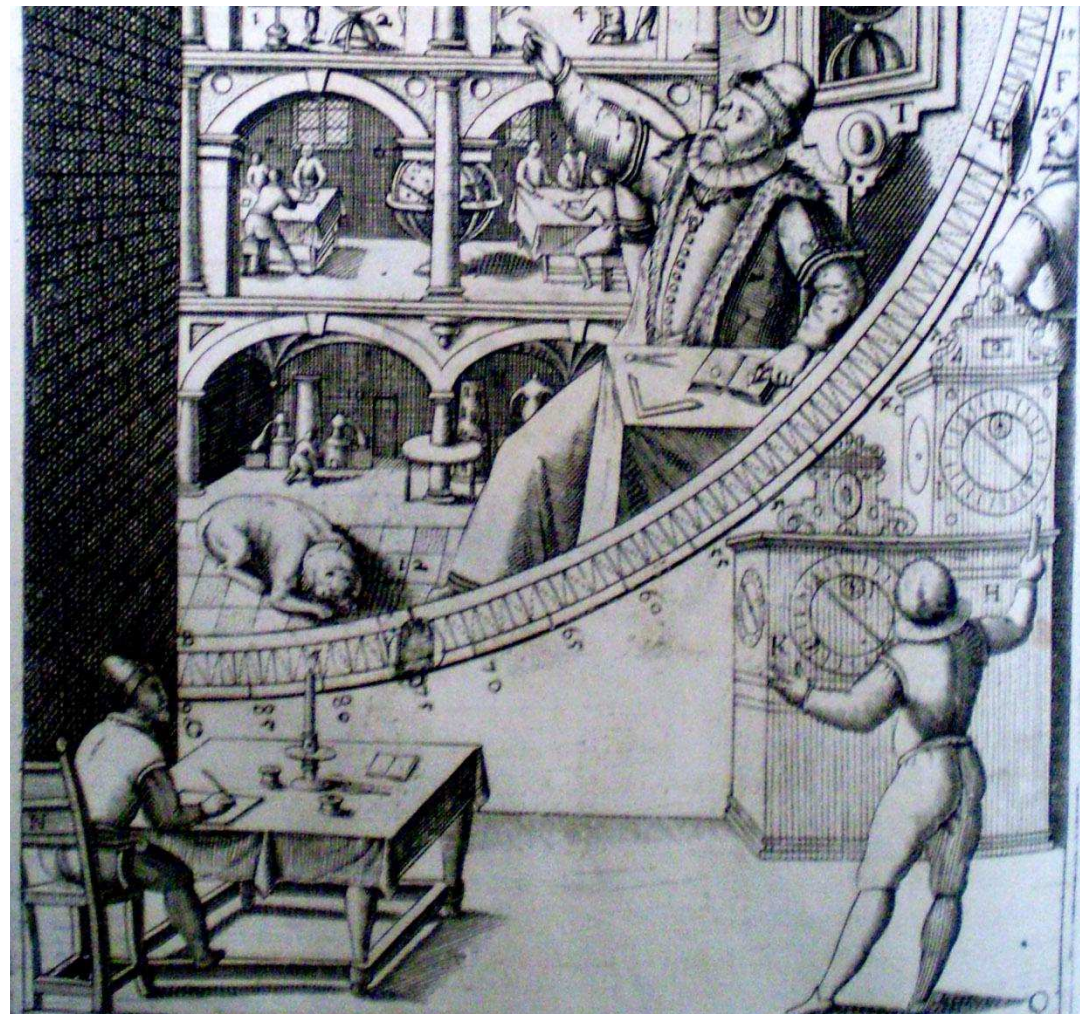
(in Prag seit 1599)



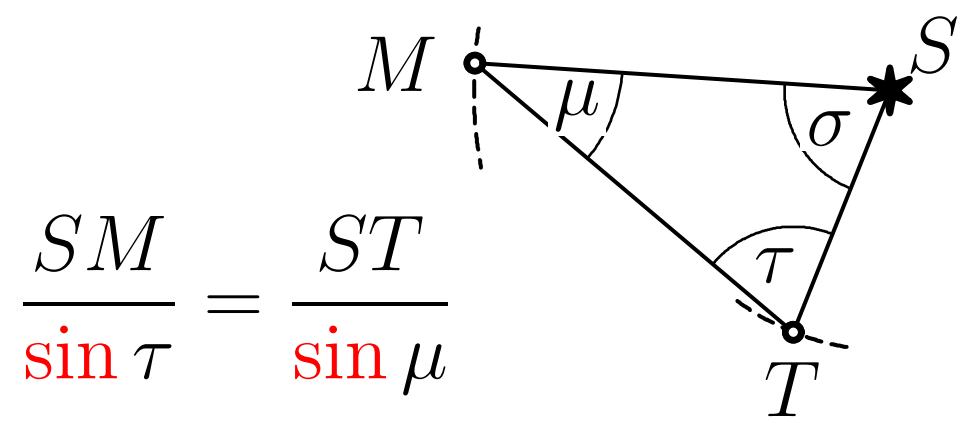
(in Prag seit 1600)



(Abb. von Kepler 1609)

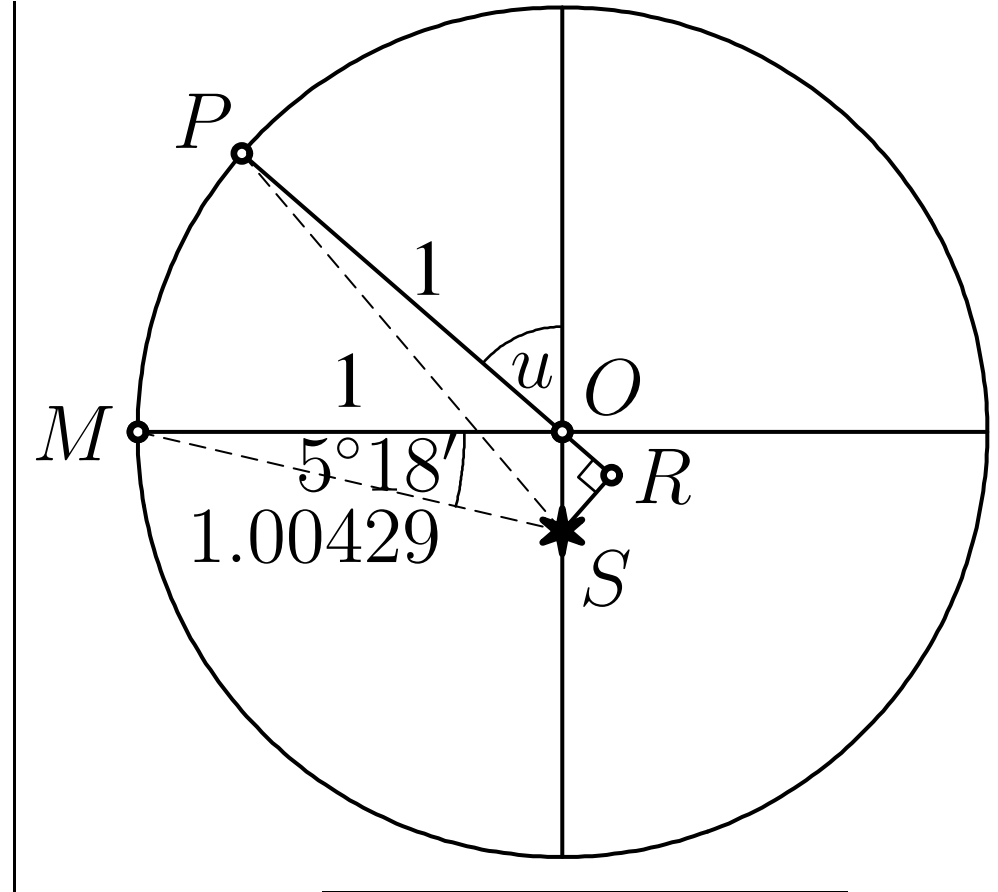
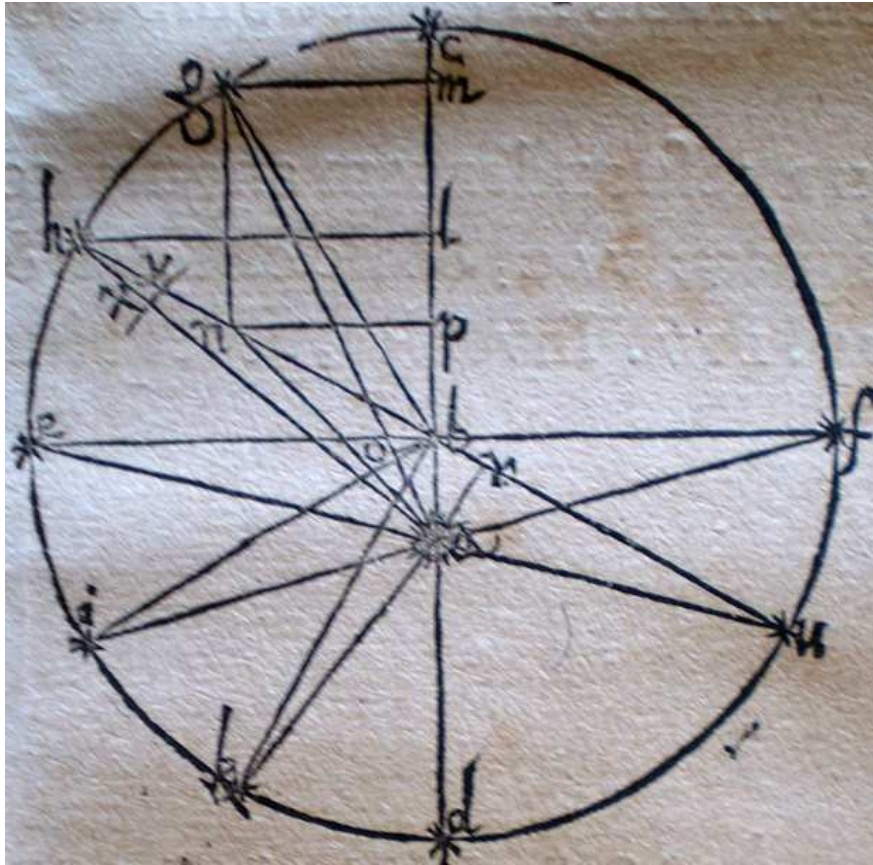


viele extrem genaue Abst.  $SM$



$$\frac{SM}{\sin \tau} = \frac{ST}{\sin \mu}$$

**Problem:** Bei **Mars** ist für exzentrisches Kreismodell  
 Abst.  $SM$  bei **Opticæ maximæ** um Faktor 1.00429 **zu gross !!**

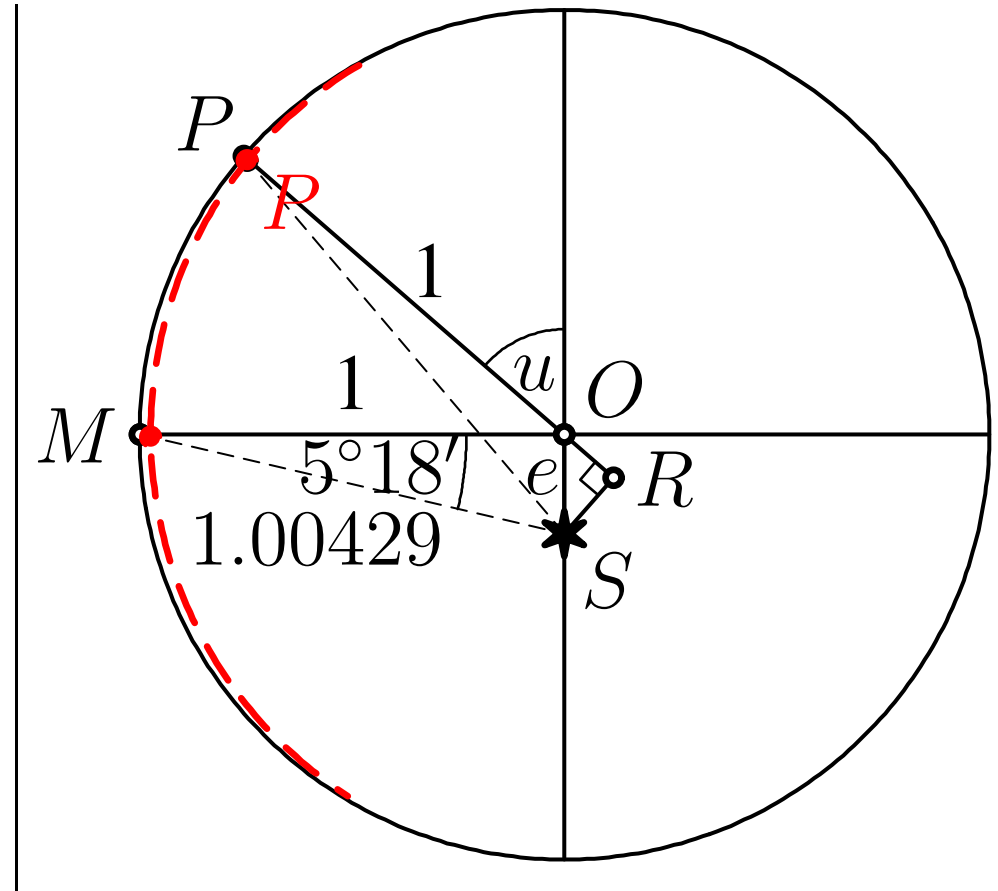
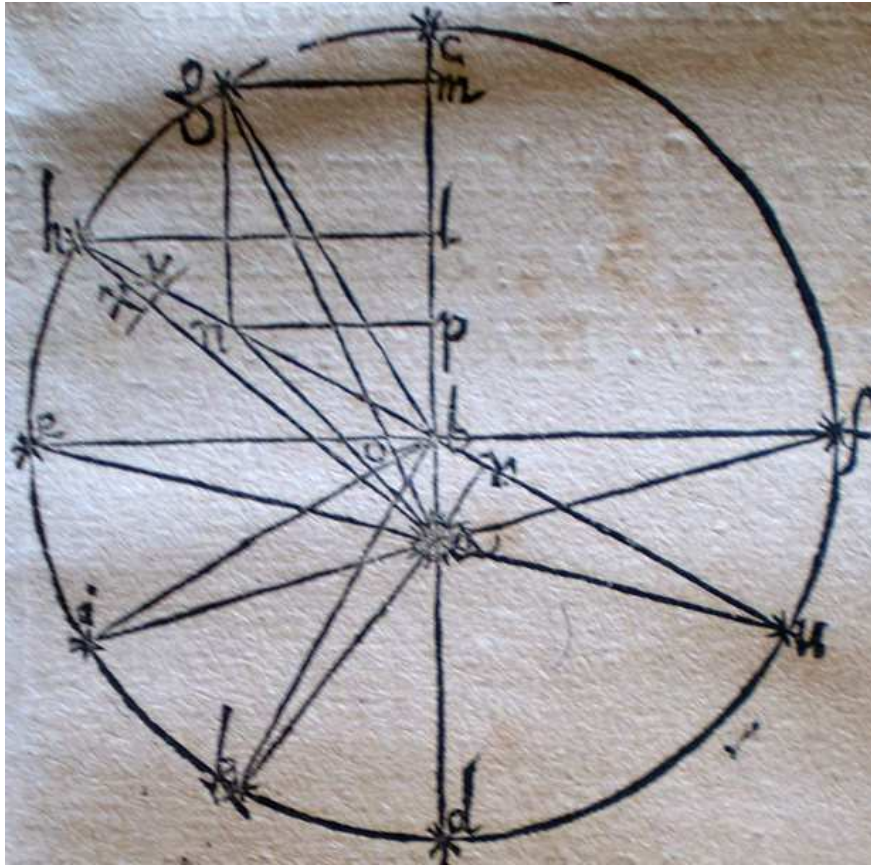


**Kepler:**  $1.00429 = 1 / \cos 5^{\circ}18'$



plane nihil dictum esse, itaque futilem fuisse meum de Marte triumphum; forte fortuito incido in secantem anguli  $\zeta$ .  $18'$ . quæ est mensura æquationis Opticæ maximæ. Quem cum viderem esse 100429, hic quasi e somno expergefactus, & novam lucem intuitus, sic cœpi ratio-

# Kepler “somno expergefactus, & novam lucem intuitus”:



Idee: **ersetze Hypotenuse durch Kathete**  $P \mapsto P$

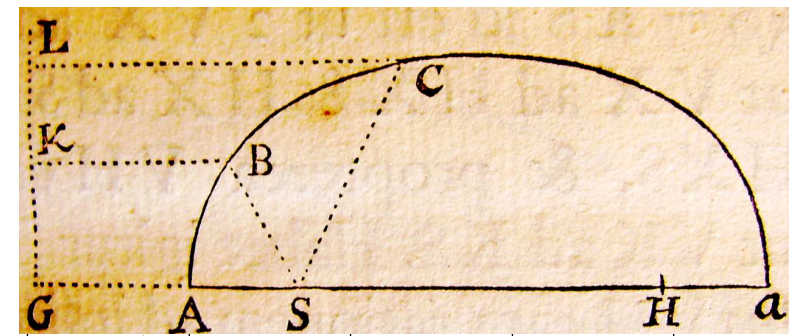
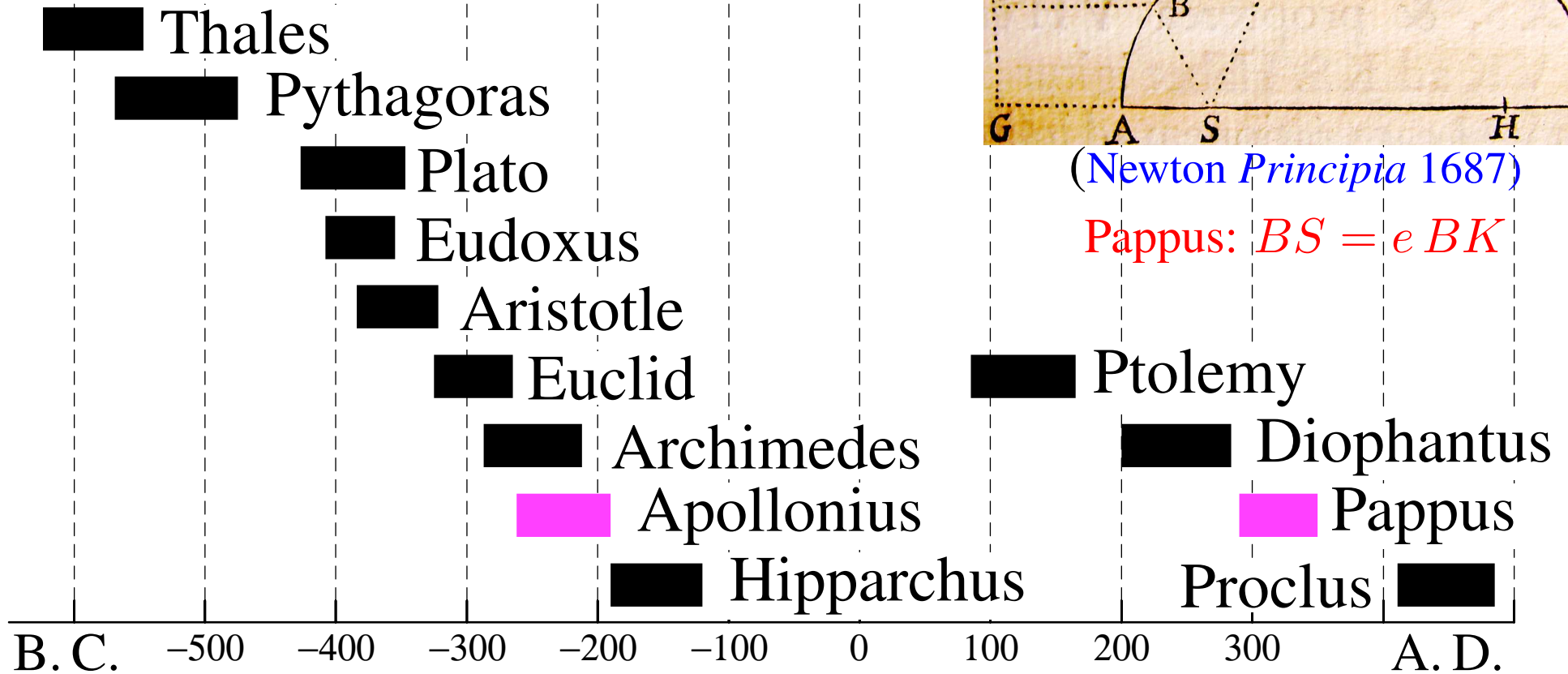
$$PS = PR = 1 + e \cos u$$

Eigentlich absurde Idee, **ABER:**

“Diese Abstände werden bestätigt durch sehr zahlreiche und sehr genaue Messungen” (Ende von Kap. 56)

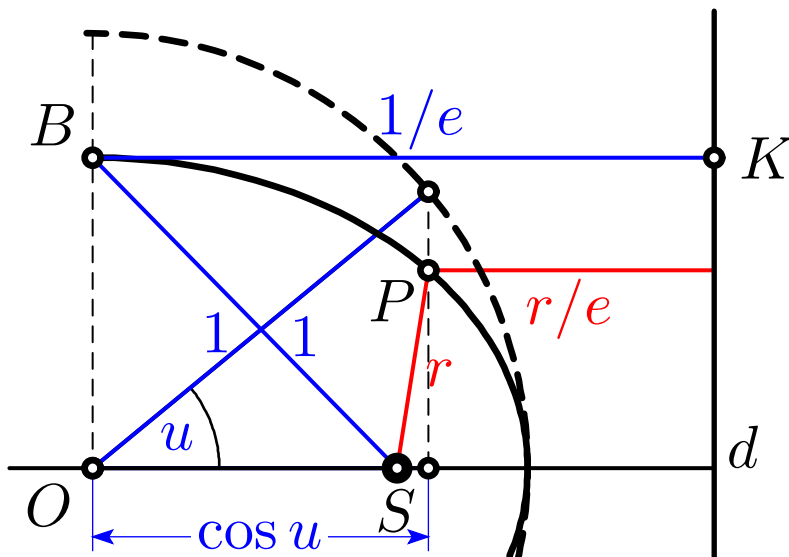


# Noch zwei Griechen ...



(Newton *Principia* 1687)

Pappus:  $BS = e BK$



$$r/e + \cos u = 1/e$$

$$r = 1 \pm e \cos u$$

$\Rightarrow$  **Bahn ist Ellipse !!**

# Johannes Kepler "de Marte triumphum" (1609):

ASTRONOMIA NOVA  
ΑΙΤΙΟΛΟΓΗΤΟΣ,  
SEV  
PHYSICA COELESTIS,  
tradita commentariis  
DE MOTIBVS STELLÆ  
**MARTIS,**  
Ex observationibus G. V.  
TYCHONIS BRAHE:  
  
Jussu & sumptibus  
**RVDOLPHI II.**  
ROMANORVM  
IMPERATORIS &c:  
  
Plurium annorum pertinaci studio  
elaborata Pragæ,  
  
*A S<sup>c</sup>. C<sup>a</sup>. M<sup>ta</sup>. S<sup>c</sup>. Mathematico*  
**JOANNE KEPLERO,**  
  
*Cum ejusdem C<sup>a</sup>. M<sup>ta</sup>. privilegio speciali*  
ANNO MDCIX Dionysianæ cō 15 c IX.

**SYNOPSIS TOTIVS OPERIS.**  
PARS PRIMA.

**PARS SECVNDA.**

**PAR TERTIA.**

**PAS QVARTA.**

**PAS QVINTA.**

# Zoom auf PARS TERTIA et QVARTA

## P A R S T E R T I A .

Secunda Epicycli seu Orbis annui. Eum vero participare etiam primo inaequalitate, hoc est, eccentricitate composita.	Stabilitur, id est, siue dependeat à Solis motu	Medio: ubi Apogrum	Res ipsa in genere. CAP. XXII.
		Apparenti: Elicitur autem rei mensura, & Apogri simul, loco Planetæ Eccentrico vel	Rei mensura. CAP. XXIII.
In usum traducitur, Constituitur	Distantiis, quarum	Res ipsa in genere. CAP. XXII.	Res ipsa. CAP. XXIV.
		Equationibus, quæ continentur ex Suppositionibus	Rei mensura. CAP. XXV.
Causæ motrices, quarum alia	in orbem circumagitur, quæ consideratur vel	Directo, observatis quaternis, CAP. XXVII.	Reciproco, ab eo quod erat demonstrandum, ad principia, observationibus quinque. CAP. XXVIII.
		Eccentricitatem causatur, Ostenditur autem	Emittens inter mobilia CAP. XXXII.
Praxis computandi ex his causis Physicis, CAP. XL.	Per se, id est, vel	Qualificatio. CAP. XXXIII.	Qualificatio. CAP. XXXIV.
		In suo fonte, qui confirmatur vel	Impedito. CAP. XXXV.
Cuiuslibet motus, CAP. XXXI.	In comparatione cum Lunæ motu, CAP. XXXVII.	Objectionum solutione de effluuiis	Libero. CAP. XXXVI.
		Quomodo insit, CAP. XXXIX.	

← Kap. 40: Flächensatz

## P A R S Q V A R T A .

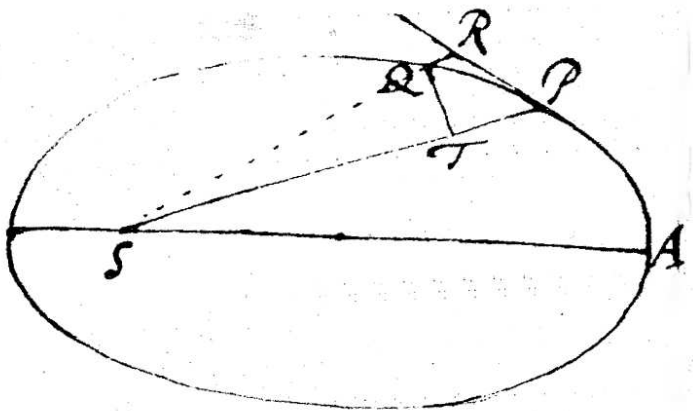
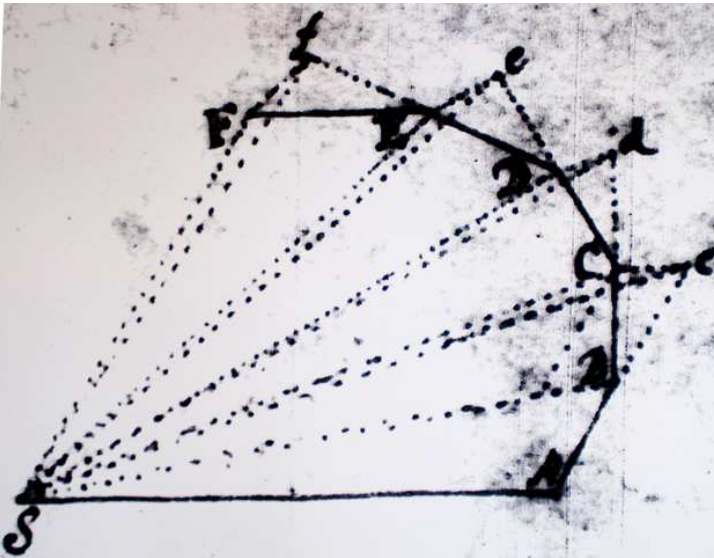
Circularis, imitatione veterum, ubi præsuppono modum causarum Physicarum, ut c. 39. Super hoc Theoria.	Principia, quæ sunt	Distantiæ, quæ inquirentur. CAP. XLII.
		Equationes, quæ ostenduntur discrepare. CAP. XLIII.
Cum præconcepit opinionem Orbitæ (quam describat Planetæ) vel	Demonstrationem ex Principiis. CAP. XLIV.	Causarum Motricum, CAP. XLV.
		Effectus Geometrici, seu figuræ decursæ. CAP. XLVI.
Ovalis angustioris ex occasione circularis nimis late deprehensæ, & certo modo causarum Physicarum atrepro, ubi Theoria	Exploratione æquationum, quæ quærentur Methodo vel	Directâ per mensurationem
		Indirectâ per inductionem sex modorum possibilium. CAP. L.
Prima Planetæ propria, ratione Eccentrici, quæ ex causis Physicis constituitur, vel	Principia, eaque vel	Effectus & circumitæ
		Efficientium principiorum. CAP. XLIX.
Possitæ falsis opinionibus, altera excedente, altera deficienti: ubi Theoria	Distantiæ, quæ	Area intra orbitam. CAP. XLVII.
		Orbitæ circa arcem. CAP. XLVIII.
Externitur demonstratis	Equationes, de quibus proximis cap. quinqu.	In usum præsentem. CAP. L.
		Eccentrici gemellæ comparata
Ad usum accommodatis	Distantiæ, quæ	Ad stabilendum Solis apparentem motum. CAP. LI.
		Plures inquirentur in selectioribus locis
Demonstrationem ex Principiis. CAP. LV.	Distantiæ, quæ	Epicycli seu Orbis Annui. CAP. LIII.
		Comparantur cum diametro Eccentrici. CAP. LIV.
Geometricæ, quæ quales quantæ sint. CAP. LVI.	Physicæ, ex correctis causis motricibus, ostensa forma motus. CAP. LVII.	
		Per æquationes iustas. CAP. LVIII.
Per æquationem & distantiarum in unam hypothesein. CAP. LIX.	Quætionum genuina demonstrata. C. LX.	

← Kap. 56: Formel der Ellipse

# Newton: Allg. Gravitationsgesetz (1684,1687)

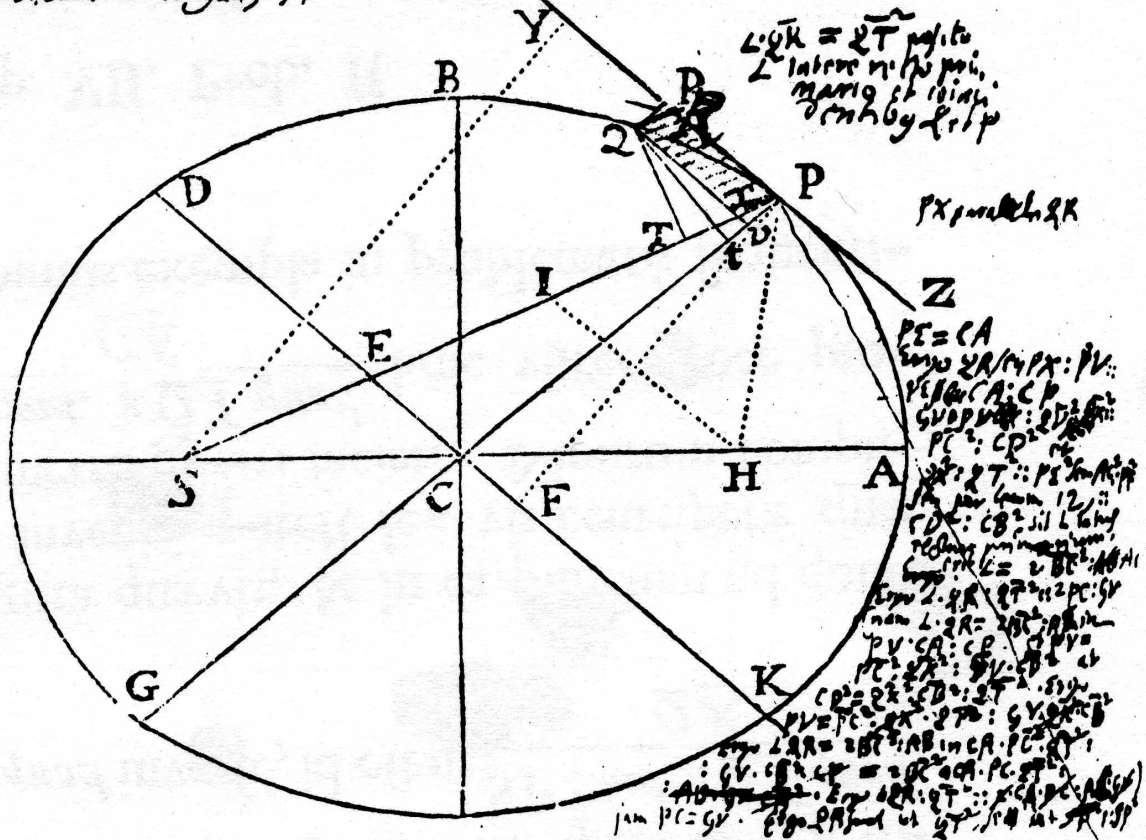
‘The Principia is .. the greatest scientific book ever written.’

(<http://mathshistory.st-andrews.ac.uk/Biographies/Newton.html>)



(from ms. Add. 3965<sup>7a</sup>)

pleatur parallelogrammum  $Q \times P R$ . Patet  $E P \propto$   
 $P F = P I + 2 I E$   $P E = P I + I E$ . *Ex*  $P E$  dimid.  $P I + I E = P I + \frac{1}{2} P I + \frac{1}{2} P I = \frac{3}{2} P I = \frac{3}{2} P H = C A$   
 nam  $P I = P H$ , quia  $D X$  *est*  $I H$  *secundum*  $U$  *in*  $U$  *perpendicularem* *ad*  
*bisectam* *anguli*,  $P F$



tem totum  $2 AC$  adæquant. Ad  $SP$  demittatur  
 (Newton's Princ. 1687, Leibniz' handwriting)

**Theorem.** Kraft prop.  $RQ$  prop.  $QT^2$  prop.  $1/r^2$ .

# Zurück zur Erde: ein Problem der Mechanik:

Voilà les fondements de la physique moderne, céleste et terrestre.

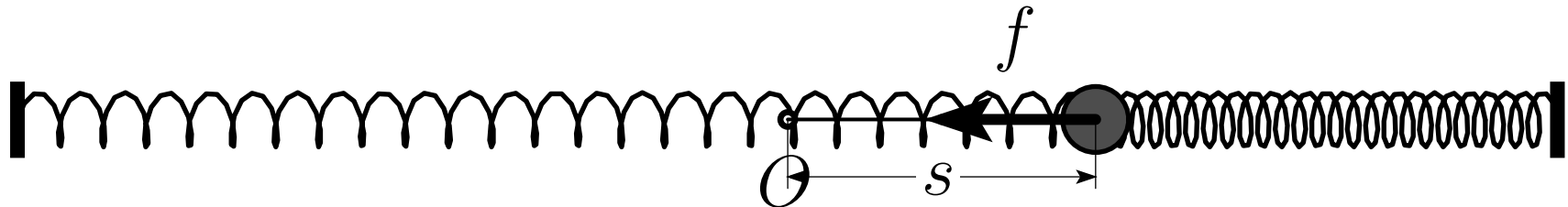
(Delambre 1821; see M. Caspar, *J. Kepler, Werke 3*, p. 256)

Wie bewegt sich ein an Federn angehängtes Gewicht ?



Bewegung verursacht durch Kraft:

*Lex 2. Mutationem motus proportionalen sp̄i si impressa*

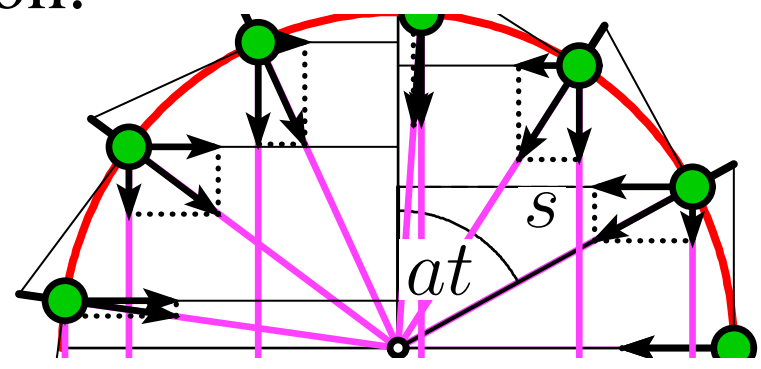
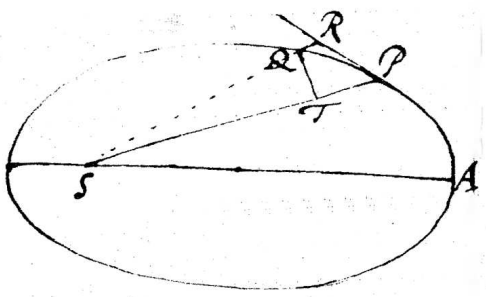


**Hookesches Gesetz** (Robert Hooke 1676):

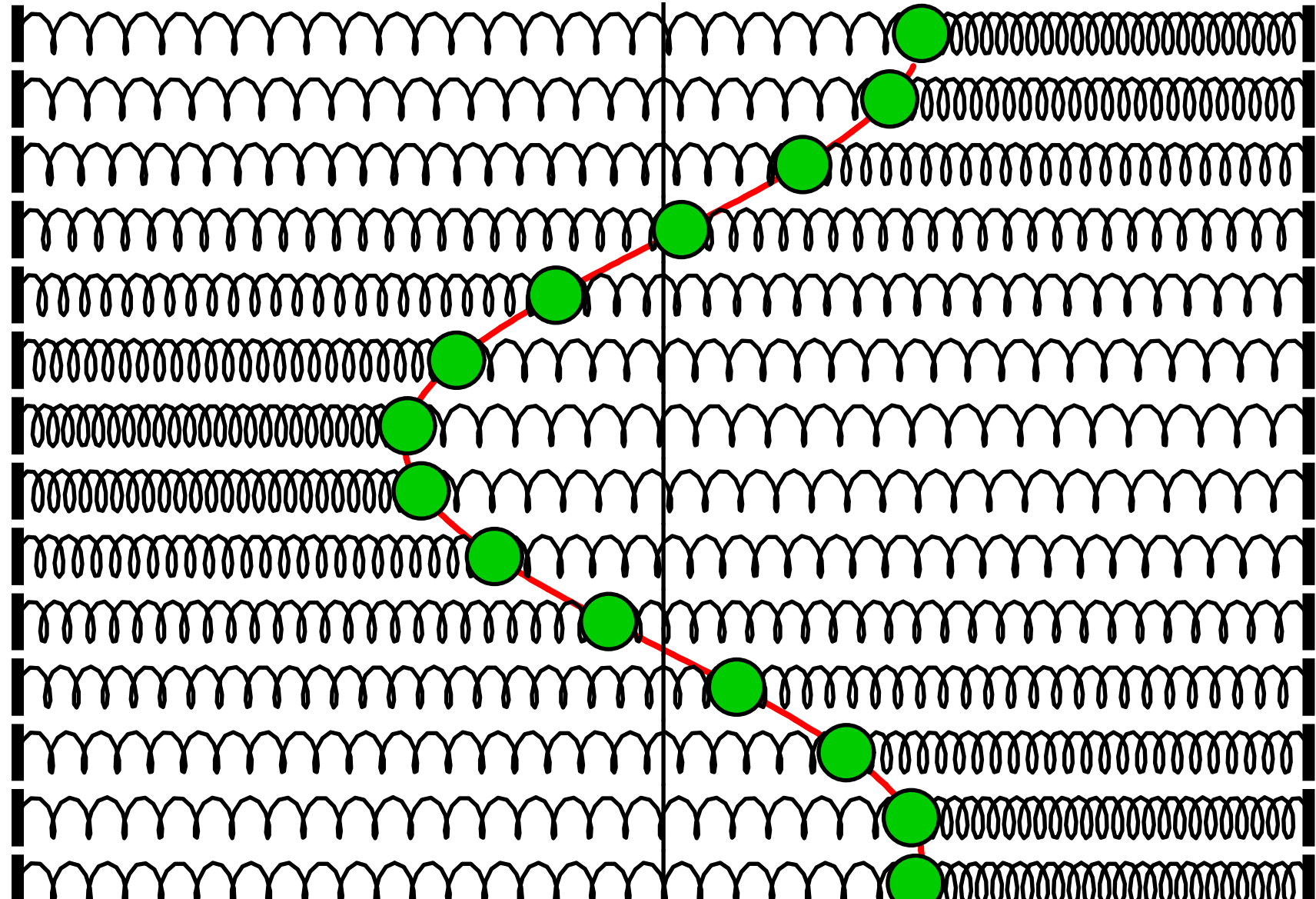
*Kraft ist proportional der Auslenkung  $f = \text{Const.} \cdot s$*

“The power of any springy body  
is in the same proportion with the extension.”

# Lsg. Inspir.v.Newton:

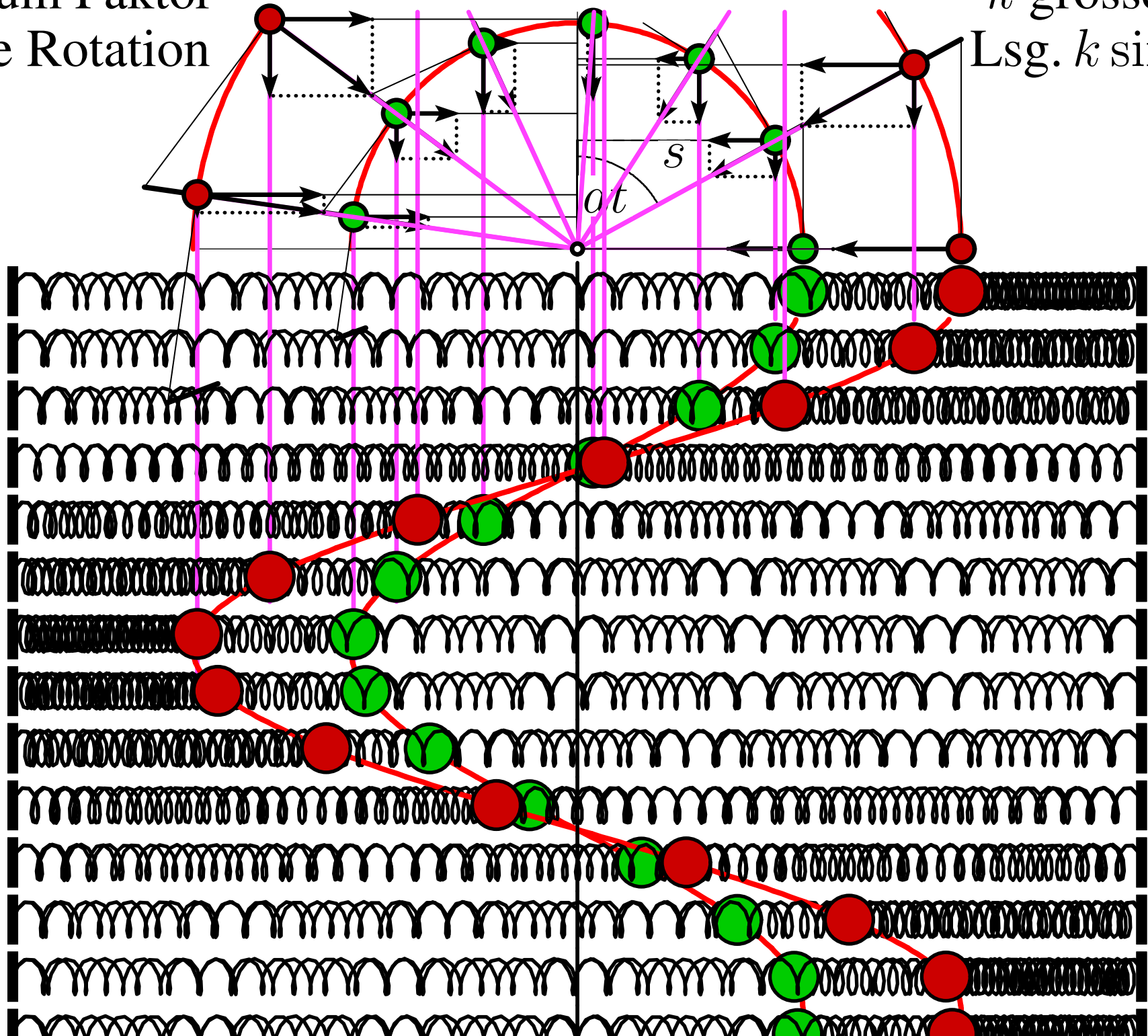


gleichf. Kreisbew.  
Horizt.kr.  $\approx \sin at$   
Hooke selbe Kraft  
 $\sin at$  ist Lösung



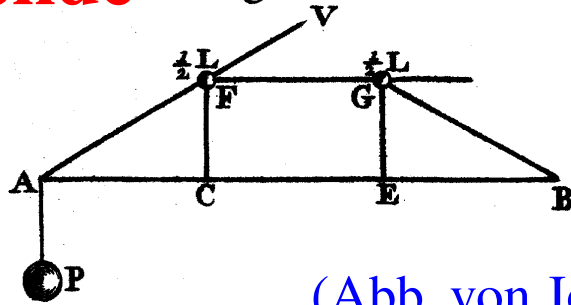
$f, s$  um Faktor  
selbe Rotation

$k$  grösser  
Lsg.  $k \sin at$



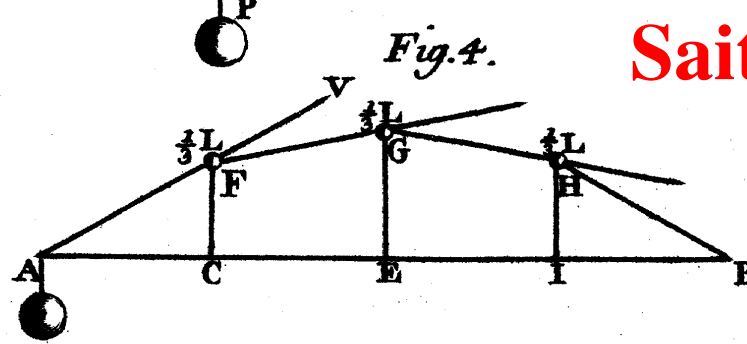
# Schwingende

Fig.3.

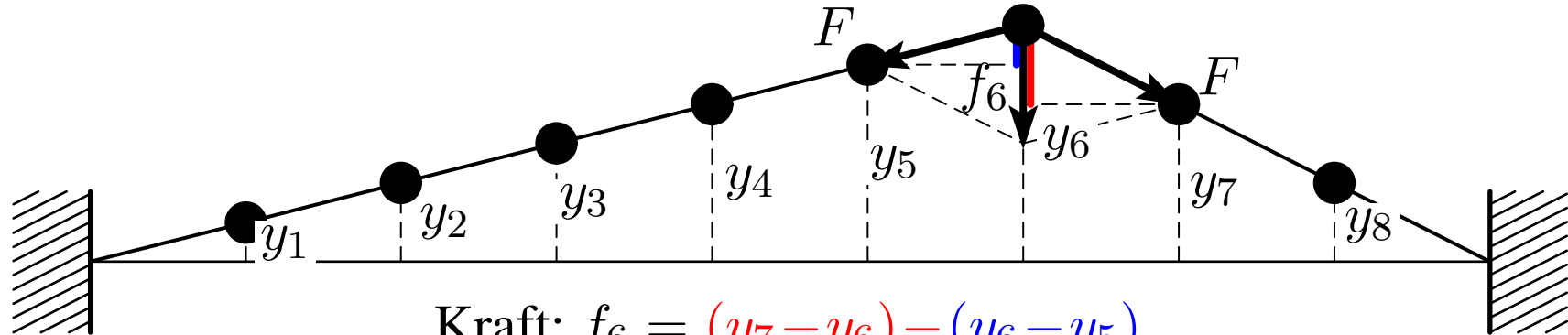


# Saite

Fig.4.



(Abb. von Johann Bernoulli 1728)

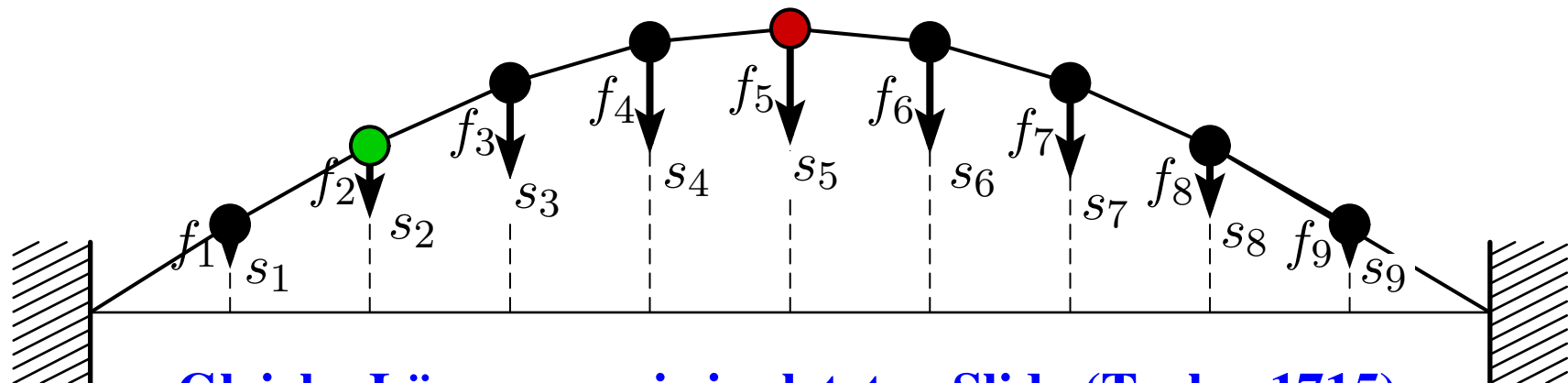


$$\text{Kraft: } f_6 = (y_7 - y_6) - (y_6 - y_5)$$

Bürgi 1:  $\sin(x + \delta) - \sin(x - \delta) = 2 \cos x \sin \delta$

Bürgi 2:  $\cos(x + \delta) - \cos(x - \delta) = -2 \sin x \sin \delta$

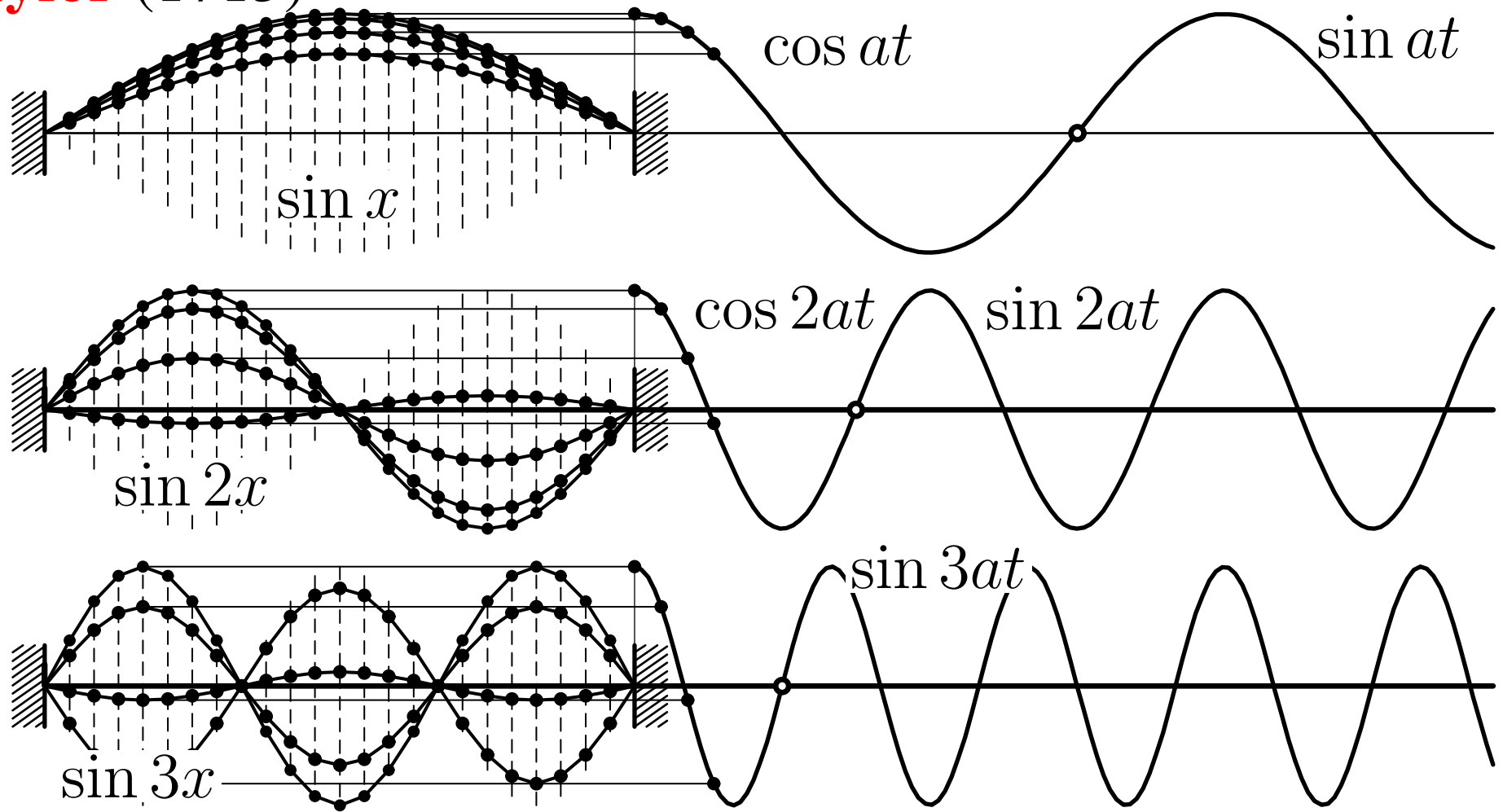
d.h., wenn  $y_i = s_i$  Sinuskurve, dann sind Kräfte  $f_i$  prop.  $-s_i$



Gleiche Lösungen wie im letzten Slide (Taylor 1715)



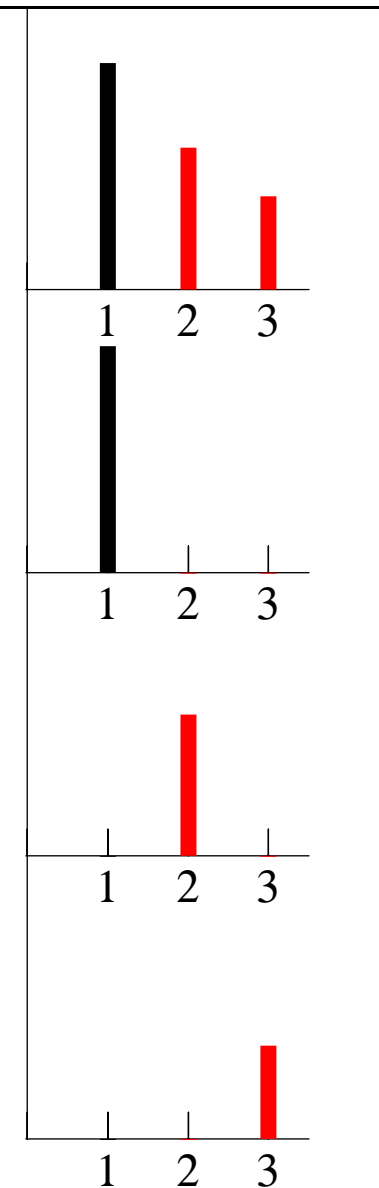
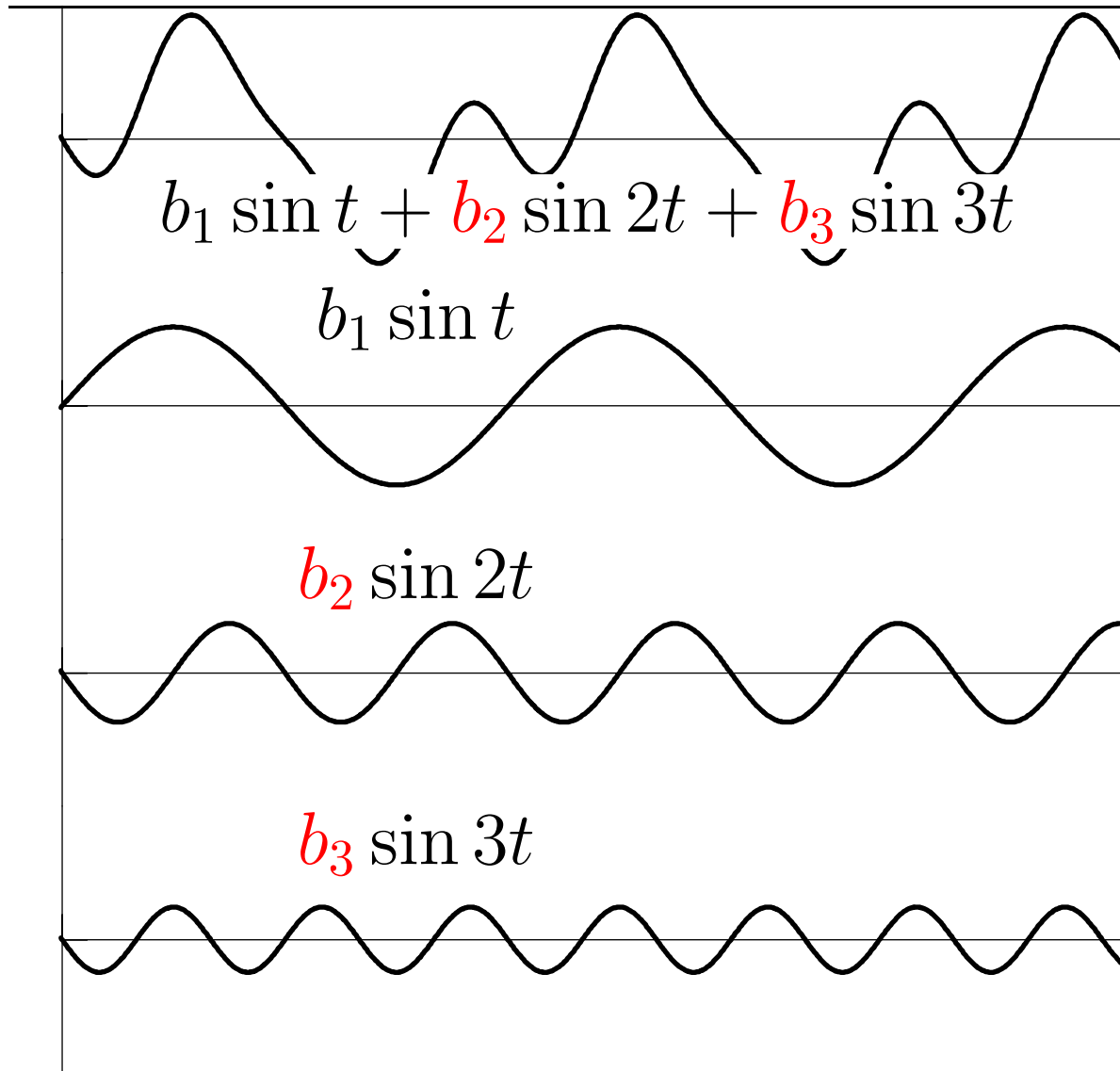
# Taylor (1715)



# Daniel Bernoulli (1753) Allg. Bew. der Saite: Überlagerung:

Grundton, harmonische O.

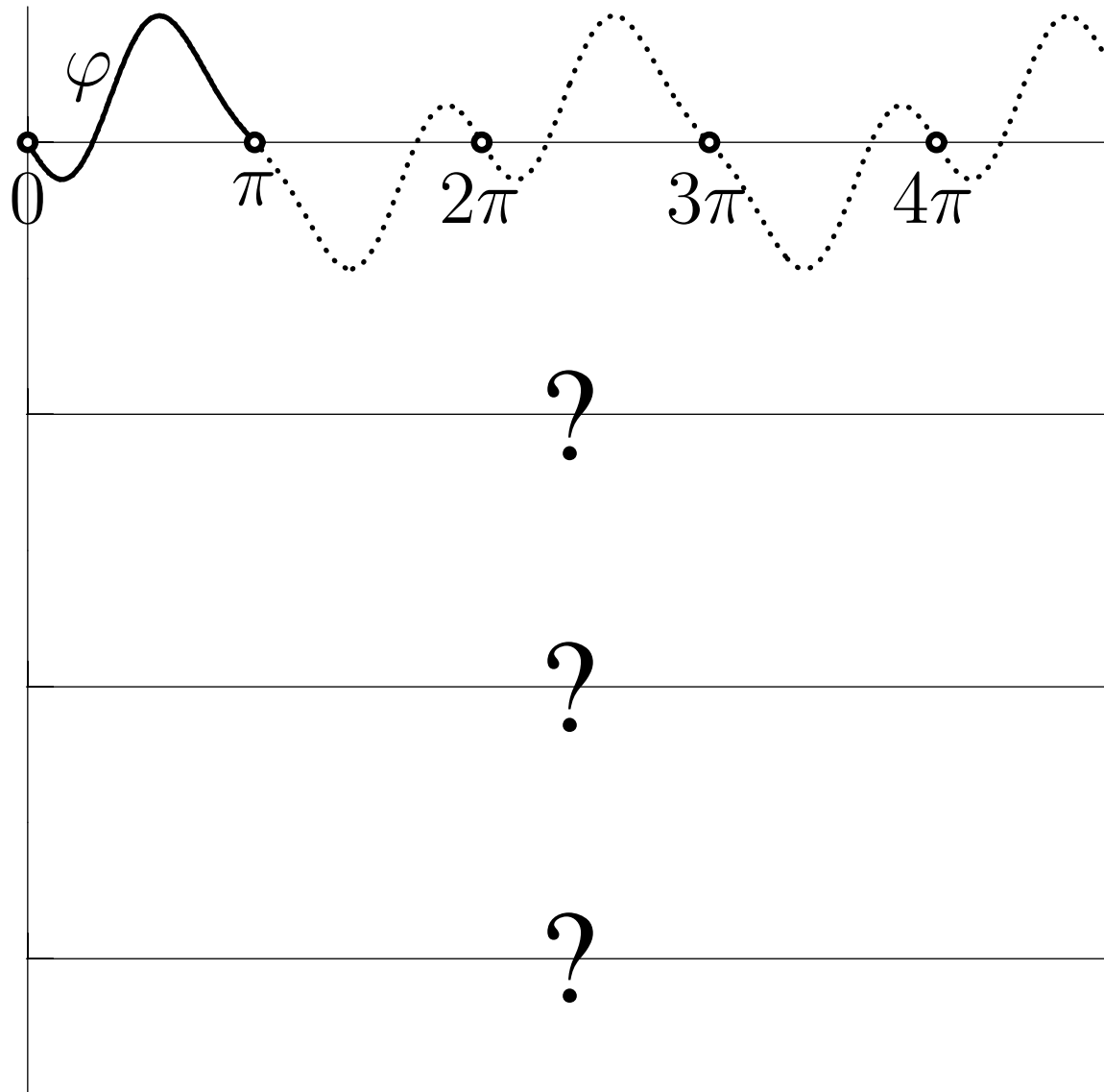
**Spektrum**



# Umgekehrtes Problem:

Funktion  $\varphi$  gegeben ( $0 \leq t \leq \pi$ ), suche  $b_1, b_2, b_3, \dots$  sodass

$$\varphi = b_1 \sin t + b_2 \sin 2t + b_3 \sin 3t + \dots$$



1	2	3
1	2	3
1	2	3
1	2	3

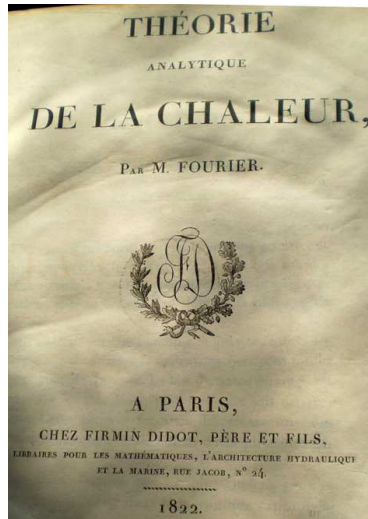
**Lösung(en): 1.** In 1764 (**E300**) schrieb Euler kühn

$$\text{nihil enim impedit,}$$
$$z = A \sin.v + B \sin.3v + C \sin.5v + D \sin.7v + E \sin.9v + \dots \text{etc.}$$

und schlug zur Berechnung von  $A, B, C, \dots$  vor, beide Seiten in die Taylorreihe bez.  $v$  zu entwickeln und (ohne Bedenken) das erhaltene unendliche lineare Gleichungssystem zu lösen.

# Idee wiederentdeckt und ausgeführt von Fourier (1807, 1822):

$$1 = a \cos y + b \cos 3y + c \cos 5y + d \cos 7y + \&c. \text{ für } -\frac{\pi}{2} \leq y \leq \frac{\pi}{2}.$$



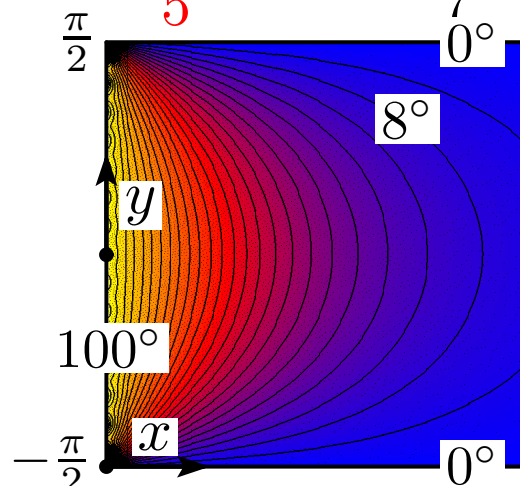
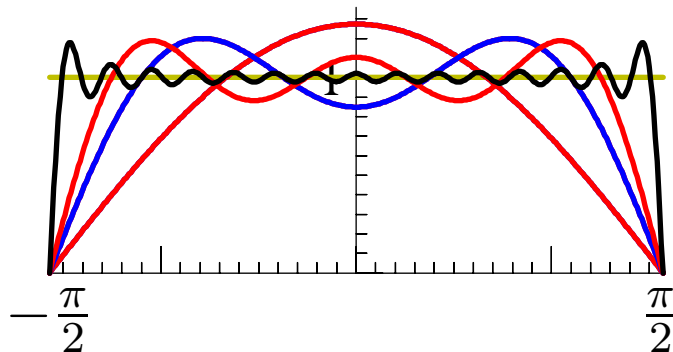
$$\begin{aligned} a + b + c + d + e + f + g &= 1 \\ a + 3^2b + 5^2c + 7^2d + 9^2e + 11^2f + 13^2g &= 0 \\ a + 3^4b + 5^4c + 7^4d + 9^4e + 11^4f + 13^4g &= 0 \\ a + 3^6b + 5^6c + 7^6d + 9^6e + 11^6f + 13^6g &= 0 \\ a + 3^8b + 5^8c + 7^8d + 9^8e + 11^8f + 13^8g &= 0 \\ a + 3^{10}b + 5^{10}c + 7^{10}d + 9^{10}e + 11^{10}f + 13^{10}g &= 0 \\ a + 3^{12}b + 5^{12}c + 7^{12}d + 9^{12}e + 11^{12}f + 13^{12}g &= 0 \end{aligned}$$

... drei Seiten Rechnung ...

$$a = \frac{3^2}{3^2 - 1} \cdot \frac{5^2}{5^2 - 1} \cdot \frac{7^2}{7^2 - 1} \cdot \frac{9^2}{9^2 - 1} \cdots = \frac{3 \cdot 3 \cdot 5 \cdot 5 \cdot 7 \cdot 7 \cdot 9 \cdot 9 \cdots}{2 \cdot 4 \cdot 4 \cdot 6 \cdot 6 \cdot 8 \cdot 8 \cdot 10 \cdots} = \frac{4}{\pi}$$

... vier Seiten Rechnung  $b, c, d, \dots$

$$1 = \frac{4}{\pi} \left[ \cos y - \frac{1}{3} \cos 3y + \frac{1}{5} \cos 5y - \frac{1}{7} \cos 7y \pm \&c. \right].$$



$$\begin{aligned} &\cos y \cdot e^{-x} \\ &\cos 3y \cdot e^{-3x} \\ &\cos 5y \cdot e^{-5x} \\ &\dots \end{aligned}$$

**Lösung 2.** 14 Jahre und 404 Arbeiten nach (**E300**), hatte Euler (**E704**) schliesslich “Nuper autem se mihi obtulit idea ...”.

Nämlich die Reihe

$$\varphi(x) = b_1 \sin x + b_2 \sin 2x + b_3 \sin 3x + \dots$$

mit, z.B.,  $\sin 2x$  multiplizieren und von 0 bis  $\pi$  integrieren

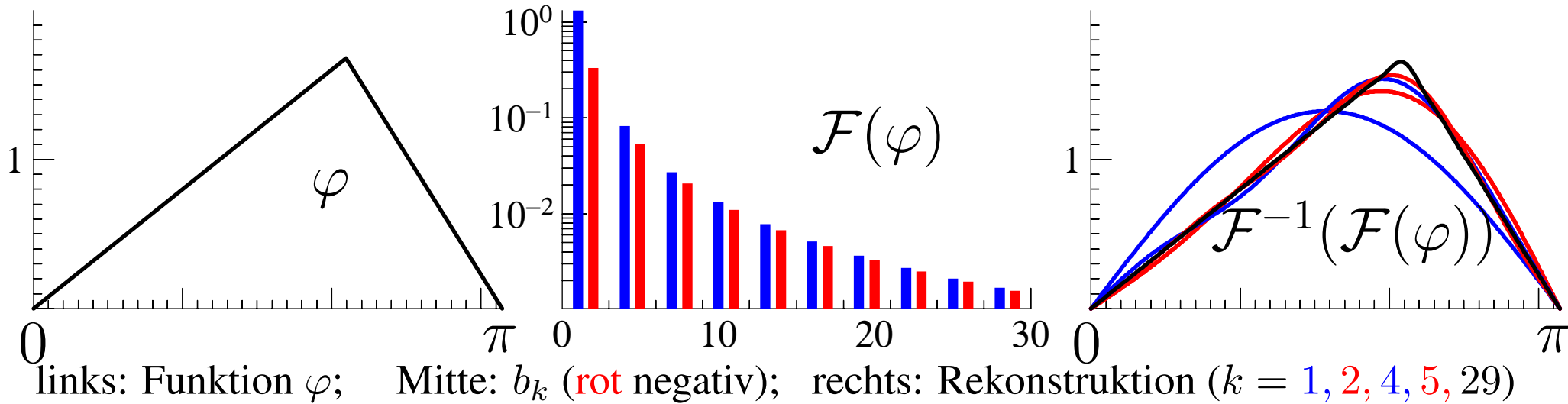
$$\int_0^\pi \varphi(x) \sin 2x \, dx = b_1 \underbrace{\int_0^\pi \sin x \sin 2x \, dx}_{=0} + b_2 \underbrace{\int_0^\pi \sin^2 2x \, dx}_{\pi/2} + b_3 \underbrace{\int_0^\pi \sin 3x \sin 2x \, dx}_{=0} + \dots$$

(fast alle Integrale sind null, “quod integrale utique evanescit...” siehe Bürgi 2), also  $b_2 = \frac{2}{\pi} \int_0^\pi \varphi(x) \sin 2x \, dx$ .

Fourier (1807, 1822) hatte unabhängig davon, nach 25 Seiten Rechnung, auch diese Idee:

**Fourier Transformation.** Für “Signal”  $\varphi(x)$ , ( $0 \leq x \leq \pi$ ), gilt für das Spektrum  $\mathcal{F}\varphi = \{b_1, b_2, b_3, \dots\}$

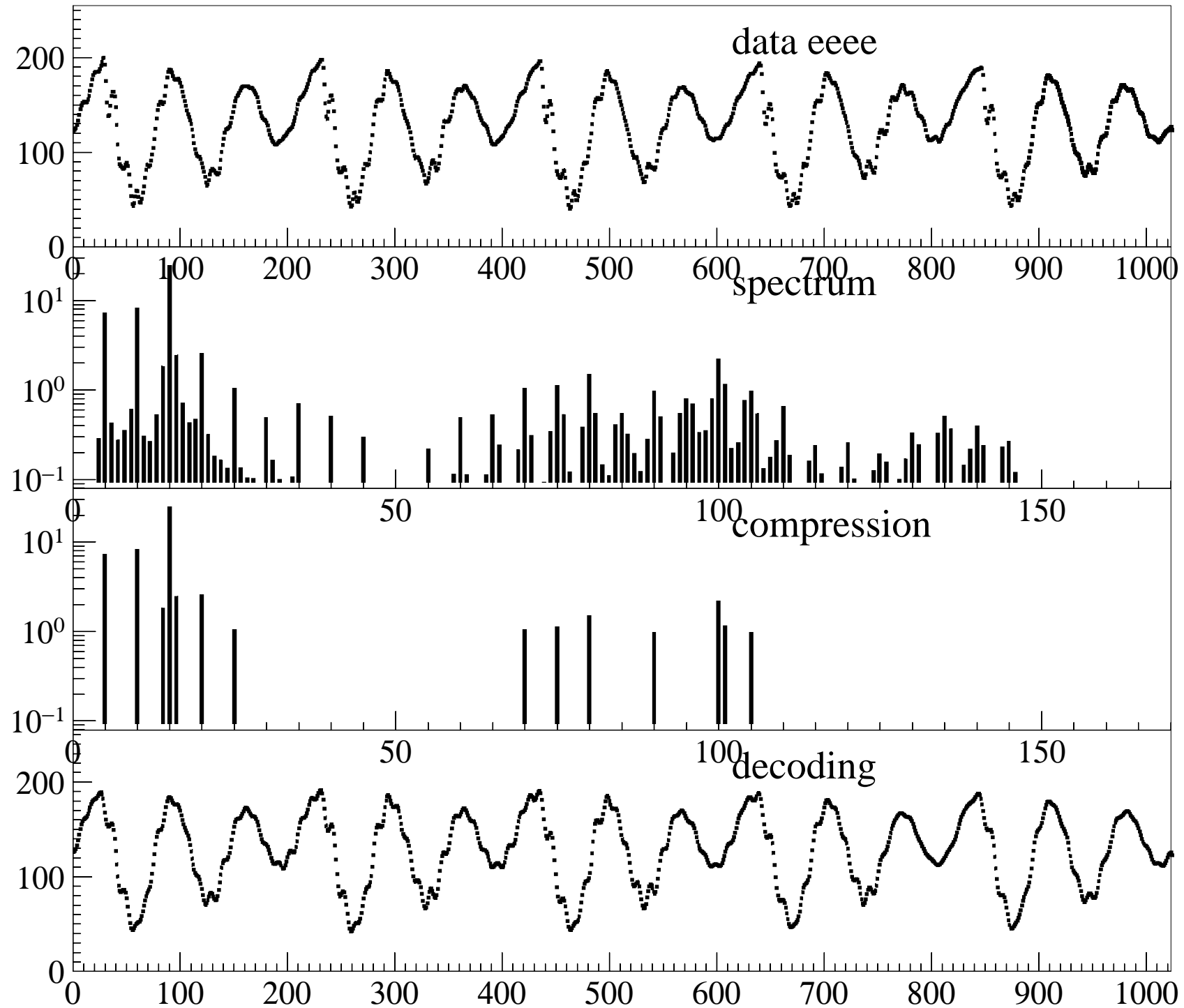
$$\mathcal{F} : b_k = \frac{2}{\pi} \int_0^\pi \varphi(x) \cdot \sin kx \, dx \quad \mathcal{F}^{-1} : \varphi(x) = \sum_{k=1}^{\infty} b_k \cdot \sin kx .$$



Das Beispiel scheint gut zu funktionieren (selbst bei einem “geknickten”  $\varphi$ ), trotzdem gab es für das 19. Jhr. einiges zu klären, z.B.:

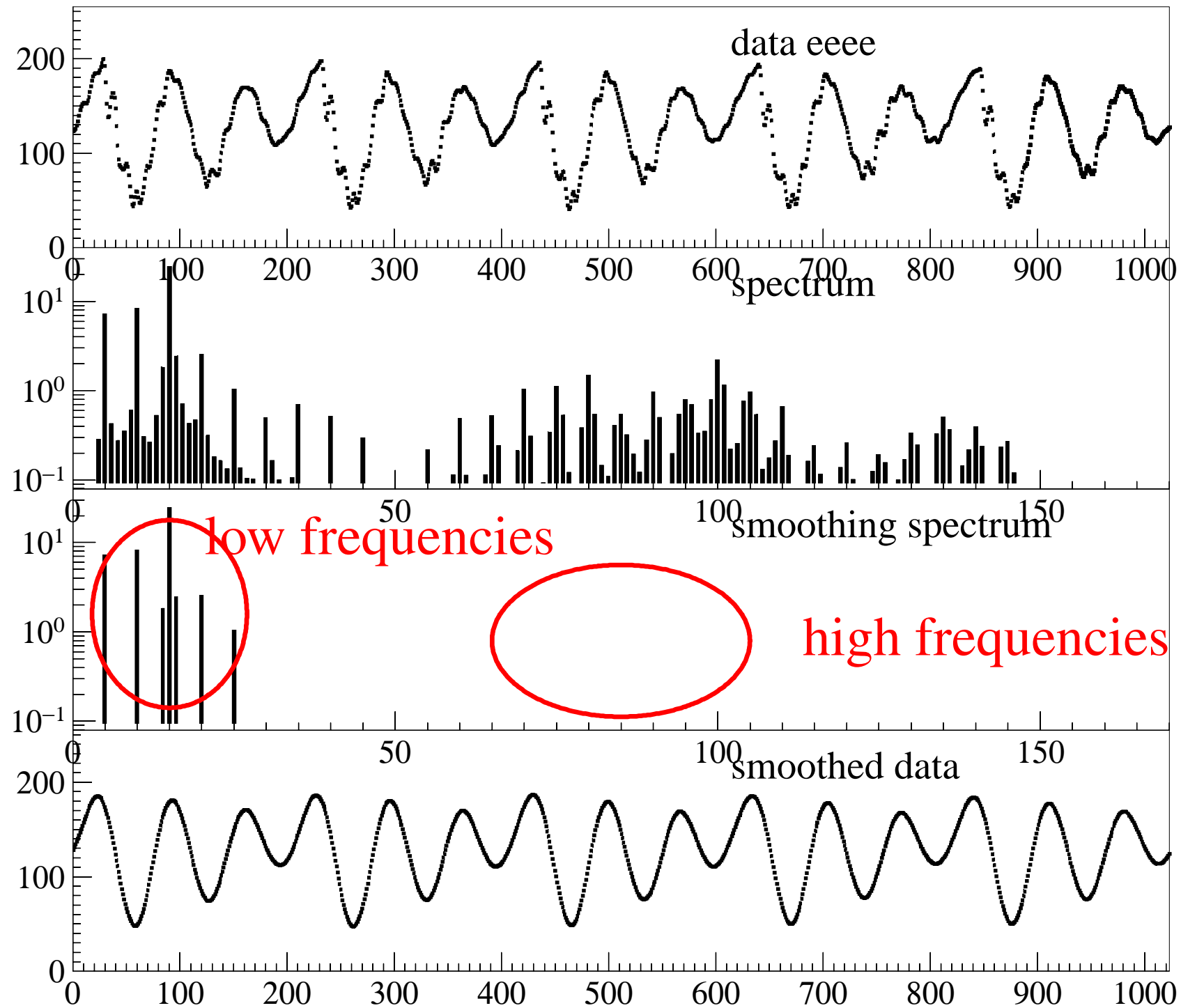
- Durfte man die unendliche Reihe gliedweise integrieren (Weierstrass)
- Was ist eigentlich ein “Integral” für die Berechnung von  $b_k$  (Riemann)
- Konvergiert die erhaltene Reihe wirklich immer gegen  $\varphi$  (Dirichlet, Fejér)

# Principle of data compression (MP3). (Kh. Brandenburg)





# Principle of data smoothing, denoising.



# Principle of JPEG data compression. (“Joint Photographic Experts Group”)



any photograph  
is cut into pieces  
of  $8 \times 8$  pixel

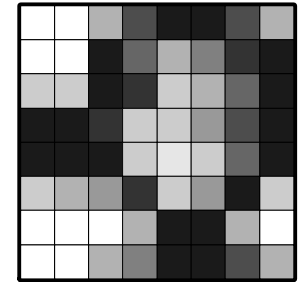
# Principle of JPEG data compression.



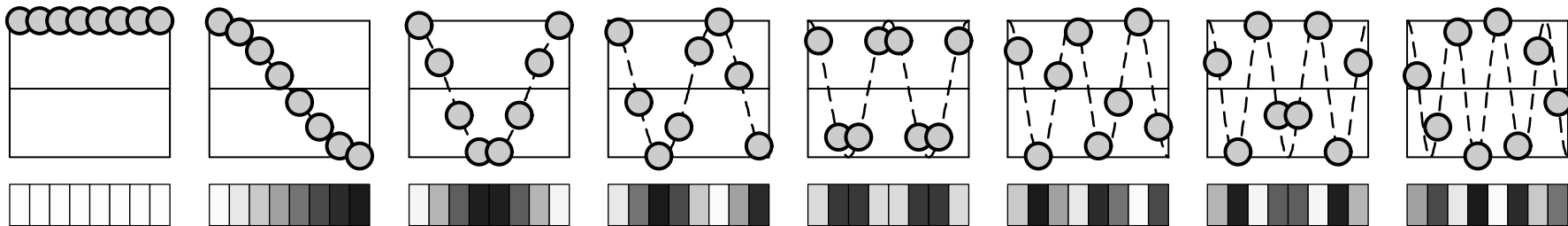
any photograph  
is cut into pieces  
of  $8 \times 8$  pixel



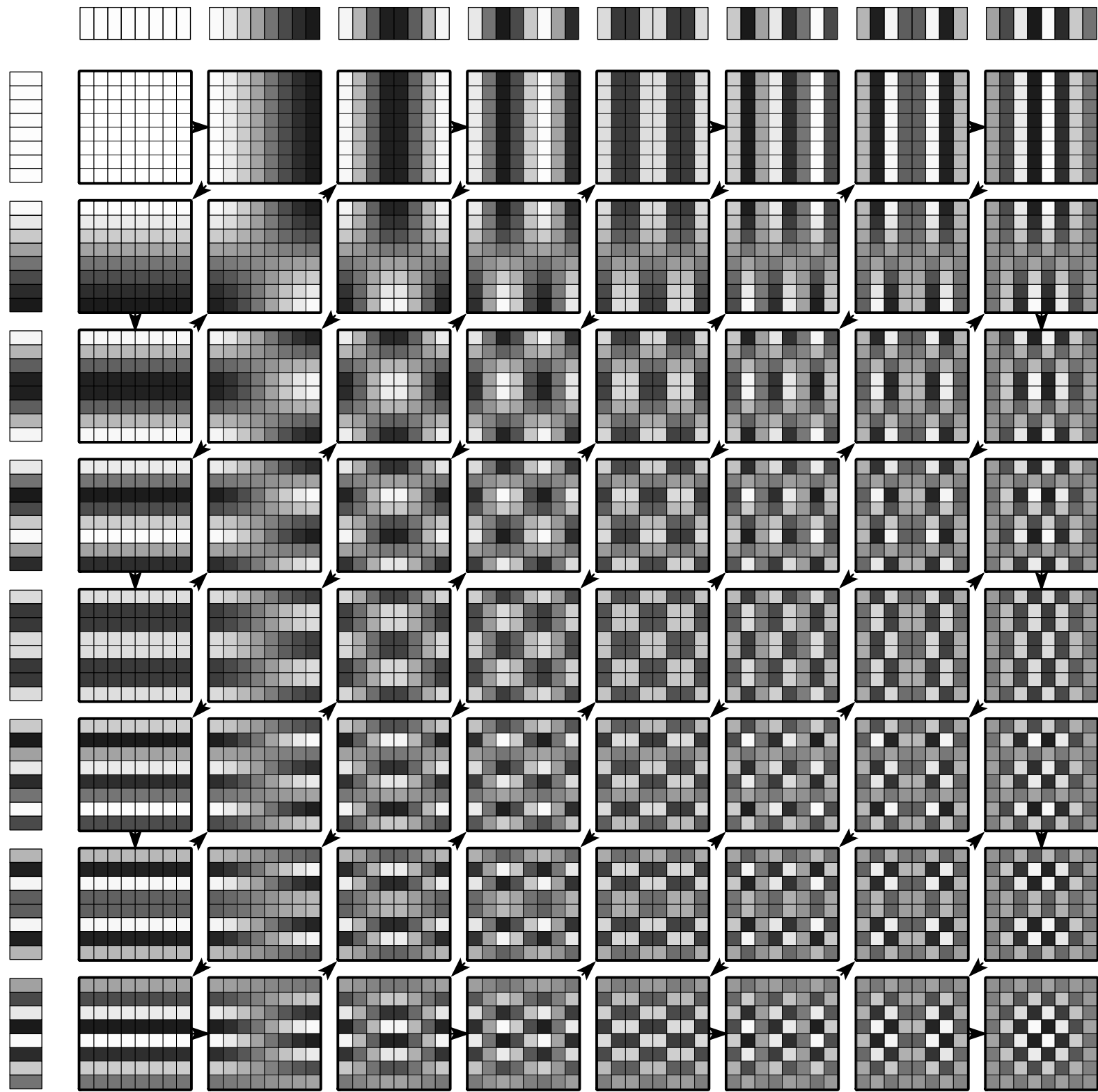
and each piece is  
treated individually



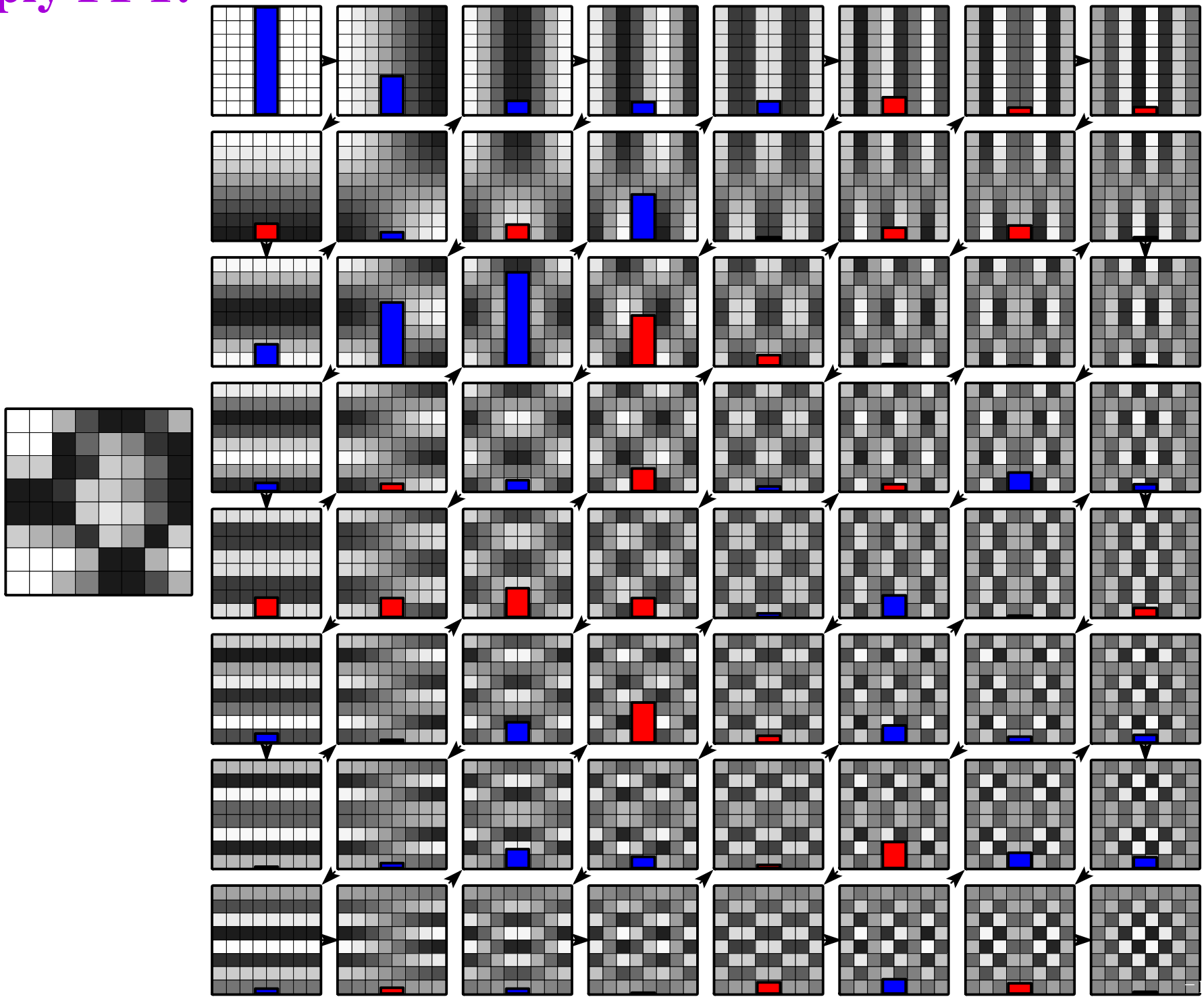
“Sound” is now two-dim.!  $\Rightarrow$  one-dim. base  $\cos k\pi x$



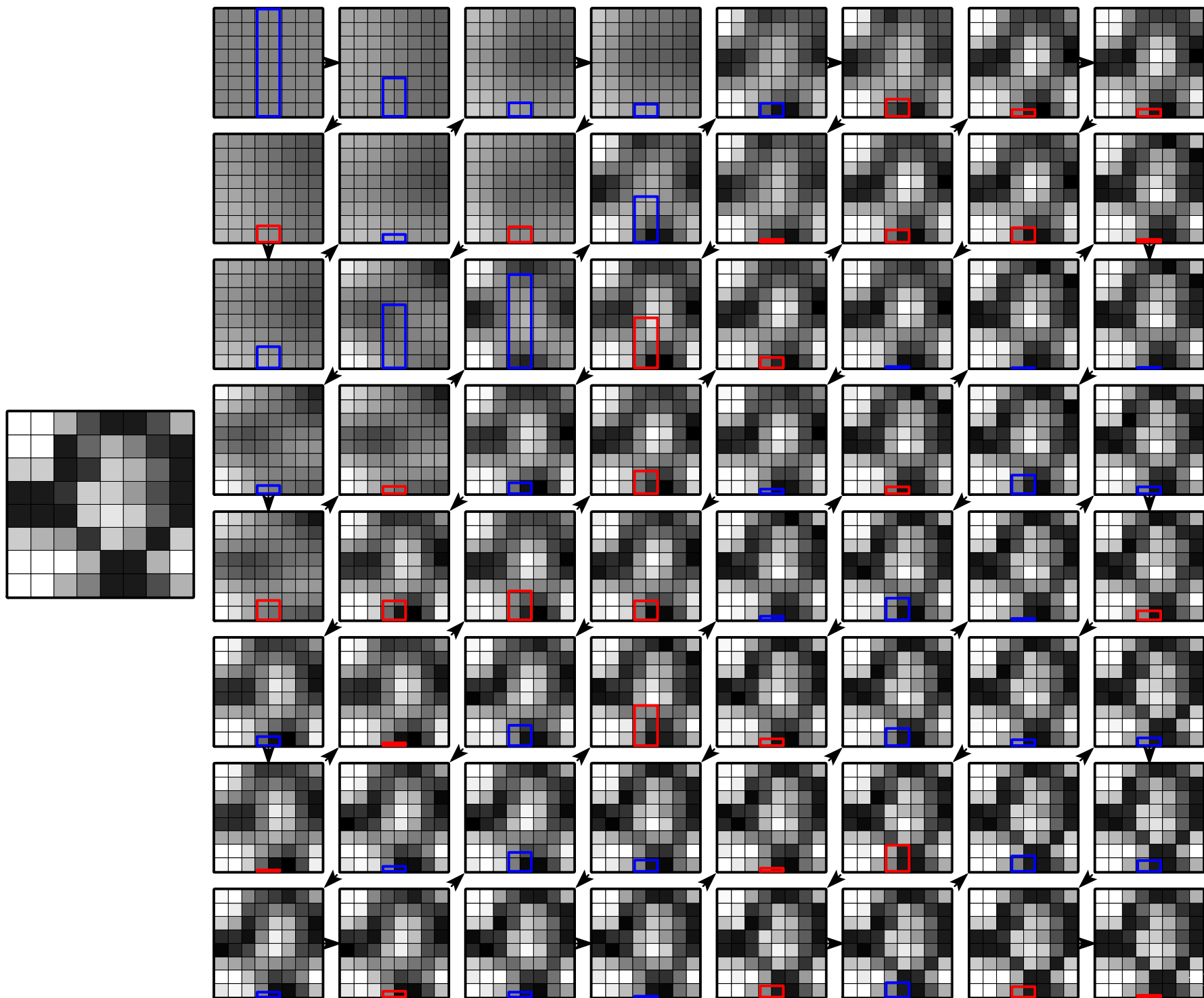
transform into two-dimensional base...  $\cos k\pi x \cdot \cos \ell\pi y \Rightarrow$



# Apply FFT:



# stages of reconstruction of image :



- Adaptation à des films et DVDs:

JPEG  $\Rightarrow$  MovingPictureExpertsGroup  $\Rightarrow$  MP3

- CDs encode sound as linear PulseCodeModulation, hence, not compressed [1981 technology] !
- Transferring CD tracks to something like an Apple iPod (portable music device) will compress the results via FFT ;
- Audio tracks on DVDs use a similar process, compressing sound frequency data (Kyle Granger).

**Vielen Dank.**

Trugarez   Grazie tanto   Thank you   Arigatō   Merci   Kiitos  
Mange Tak   Tack så mycket   Muchas gracias   Спасибо

谢谢

**Auf Wiedersehen** am 22.02.2022 zu 200 Jahre *Théorie de la Chaleur* (Fourier).